



The Higgs Discovery and Beyond

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Boulder

The Standard Model of particle physics (SM)

With the announcement of the Higgs boson discovery on July 4, 2012, all Standard Model particles have been observed.

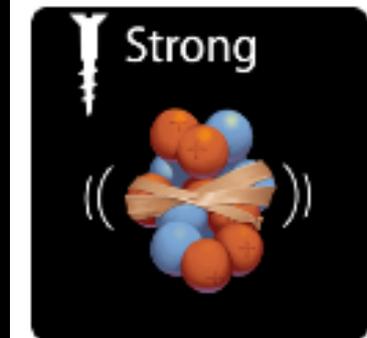
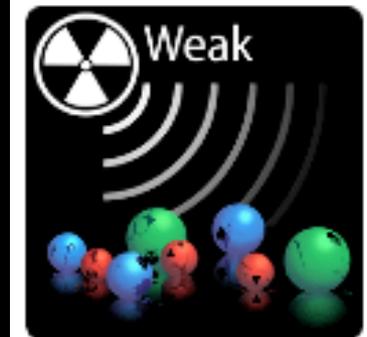
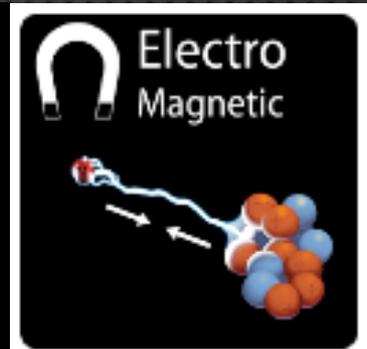
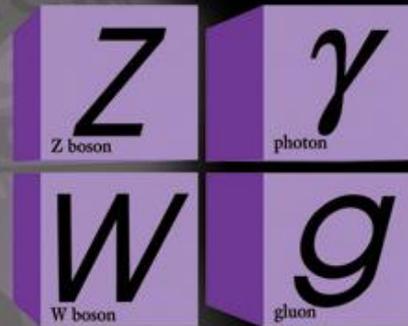
Quarks



Leptons



Forces

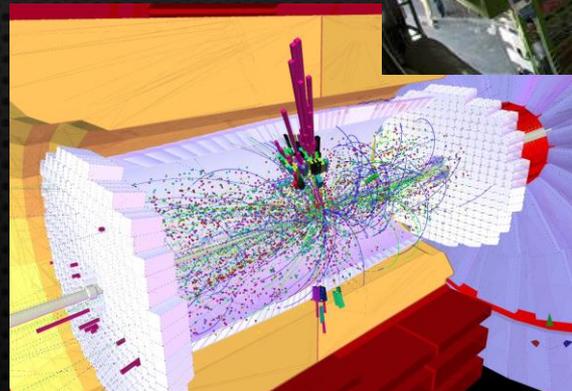
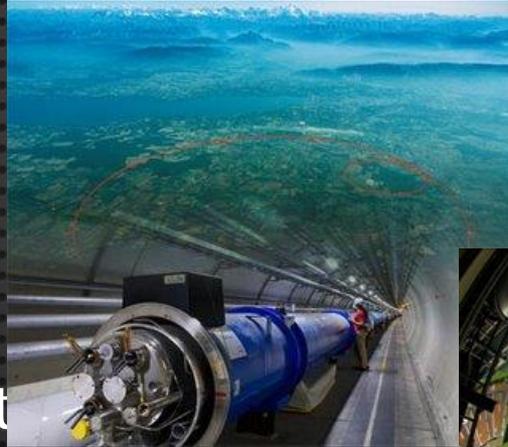


How do we get electroweak theory?

- At low energy we see E&M and weak forces
- These are unified at high energy (>1 TeV)
- The weak force contains massive force vector bosons (W^+, W^-, Z^0) but directly adding mass terms for W & Z to the theory does not work
- Use spontaneous symmetry breaking – the Higgs mechanism
- The Higgs mechanism solves two problems:
 - Mechanism to give W and Z bosons a mass in such a way as to avoid unitarity violation of WW (or ZZ) cross section at high energy
 - Also gives mass to quarks and charged leptons

How do we look for new particles?

- Need an accelerator to produce the particles.
- Need a detector to record the production and decay of the particles.
- Need to distinguish interesting events from SM events.

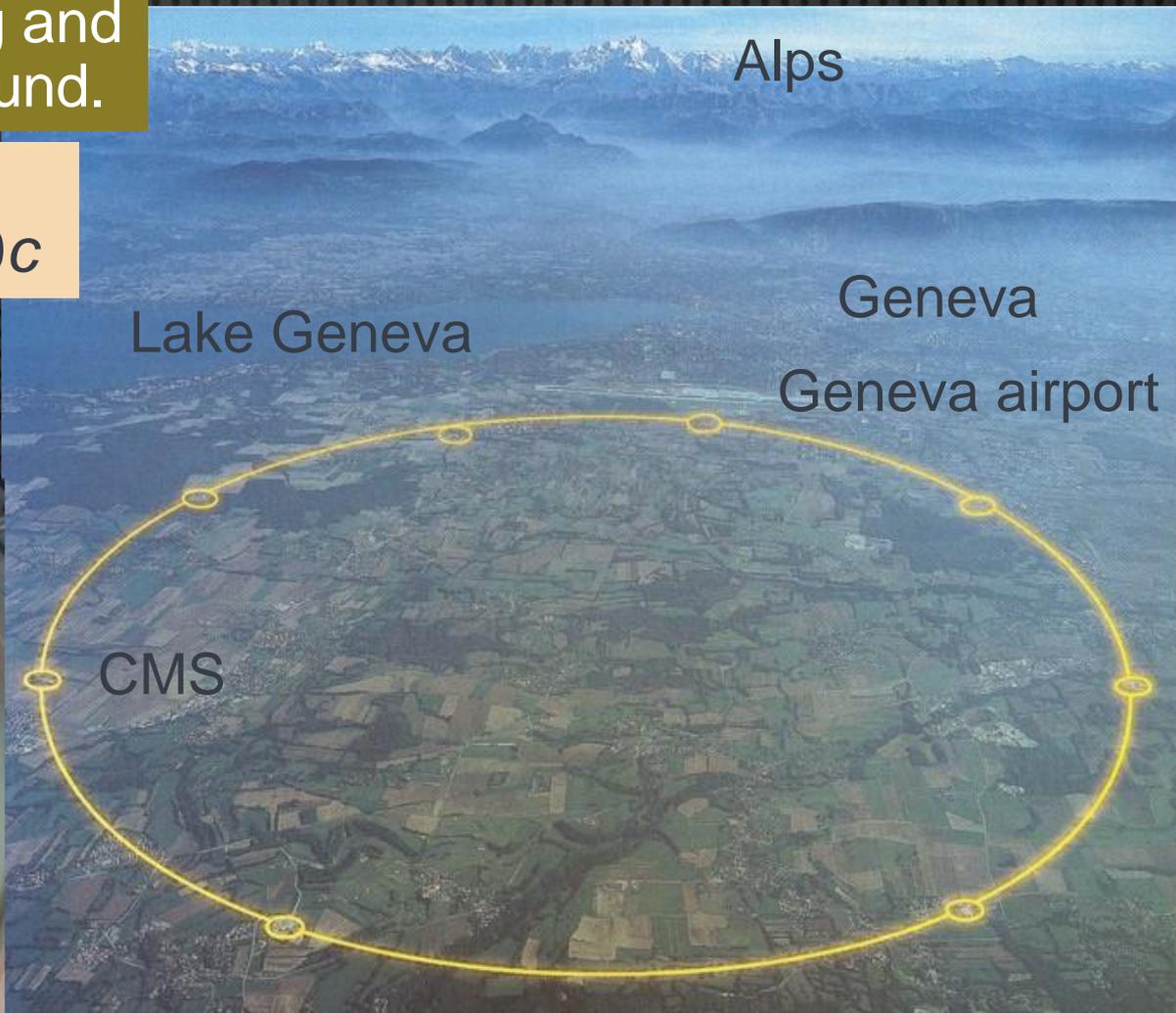


The Large Hadron Collider (LHC)

The LHC is 27 km long and 100-500 feet underground.

RF cavities accelerate protons to $0.999999999c$

8.3 T superconducting magnets keep the protons going in circles

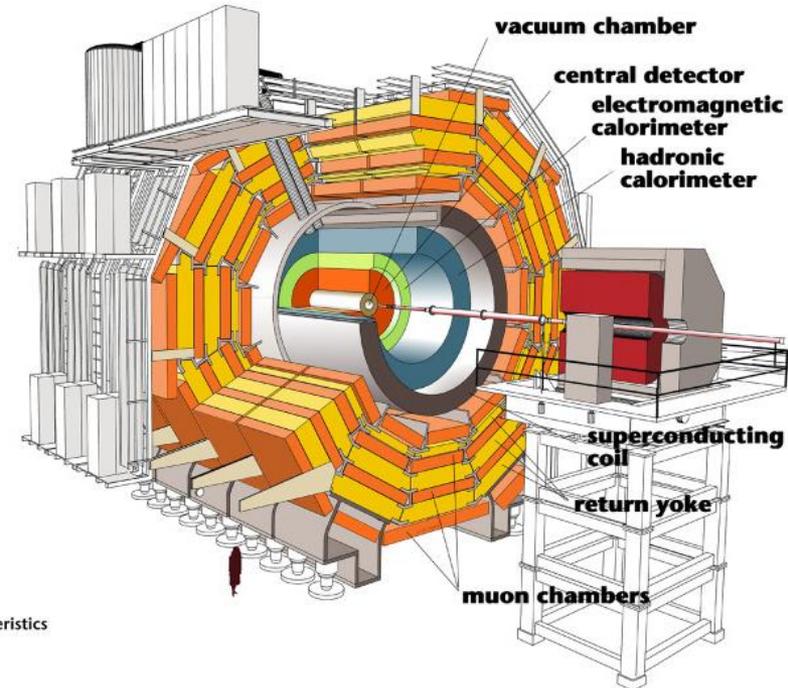
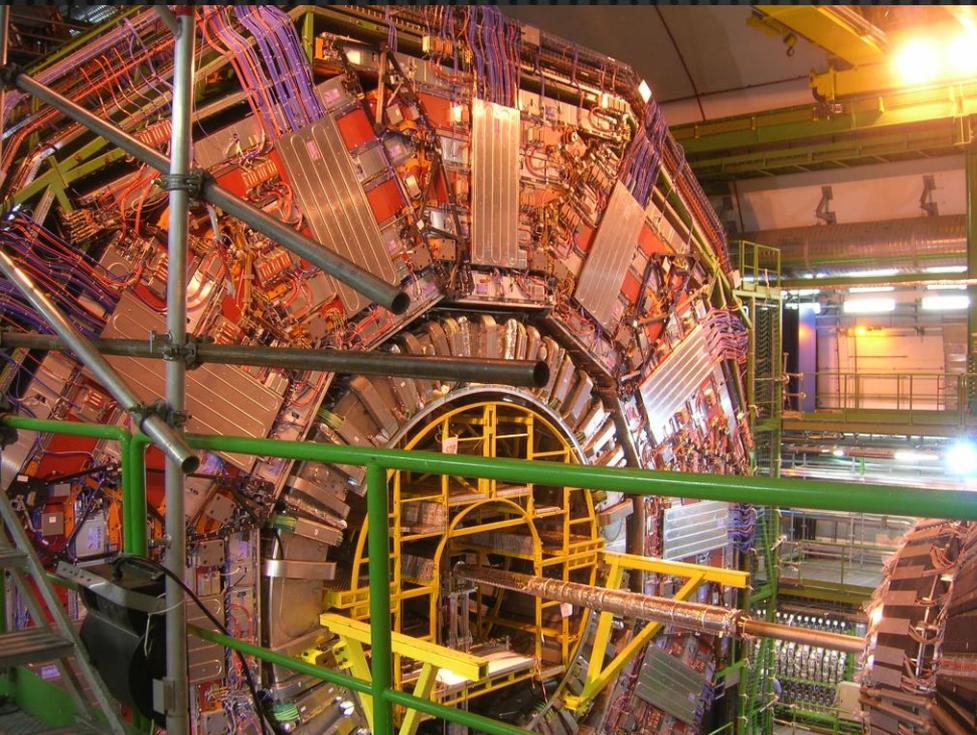


Detecting the particles

We collide protons at high energy to create new particles. But these particles immediately decay into other particles.

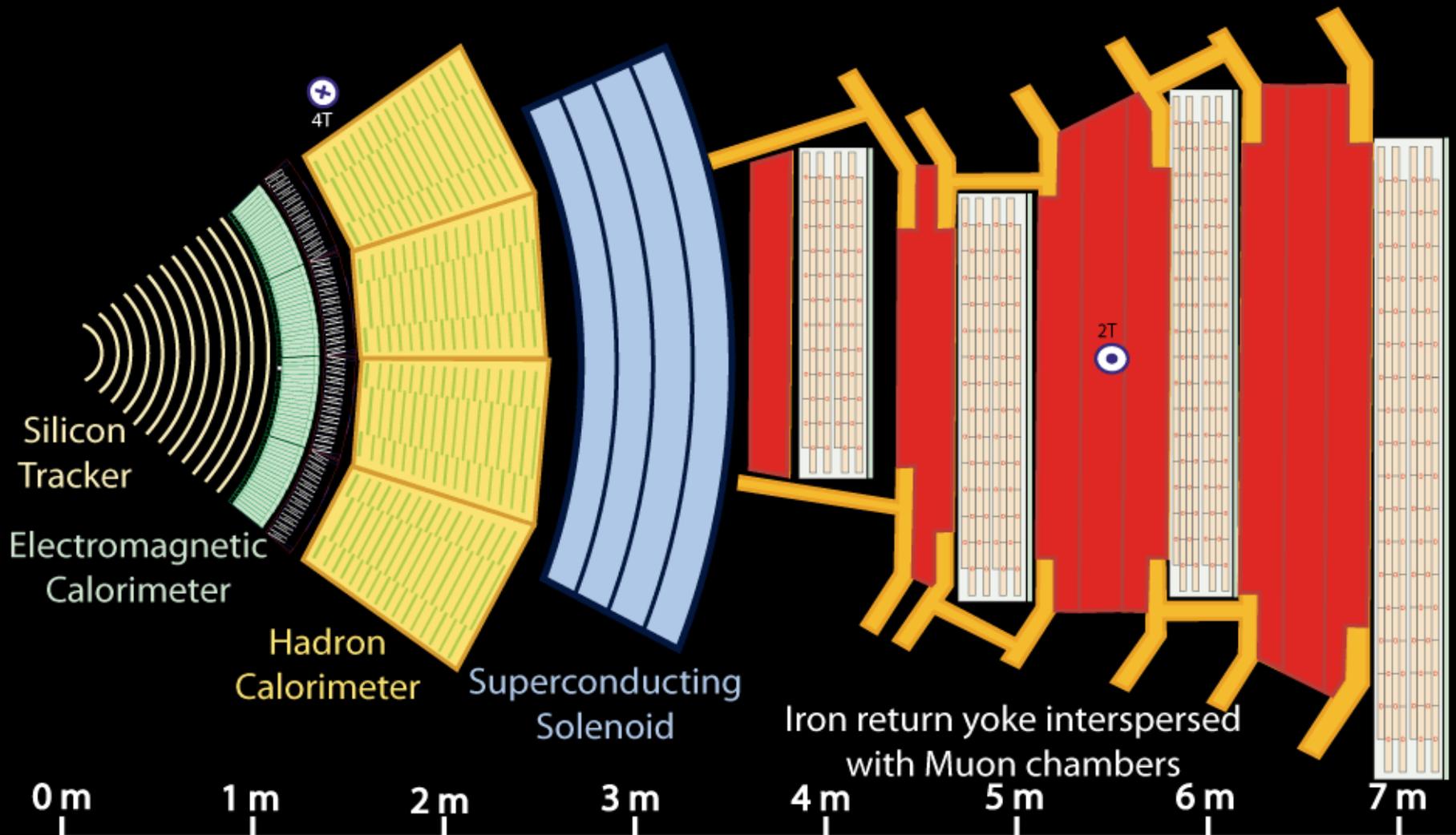
Need to detect these particles to reconstruct what happened.

The CMS detector is in a cavern 300 feet underground.



Detector characteristics

Width: 22m
Diameter: 15m
Weight: 14'500t



Key:

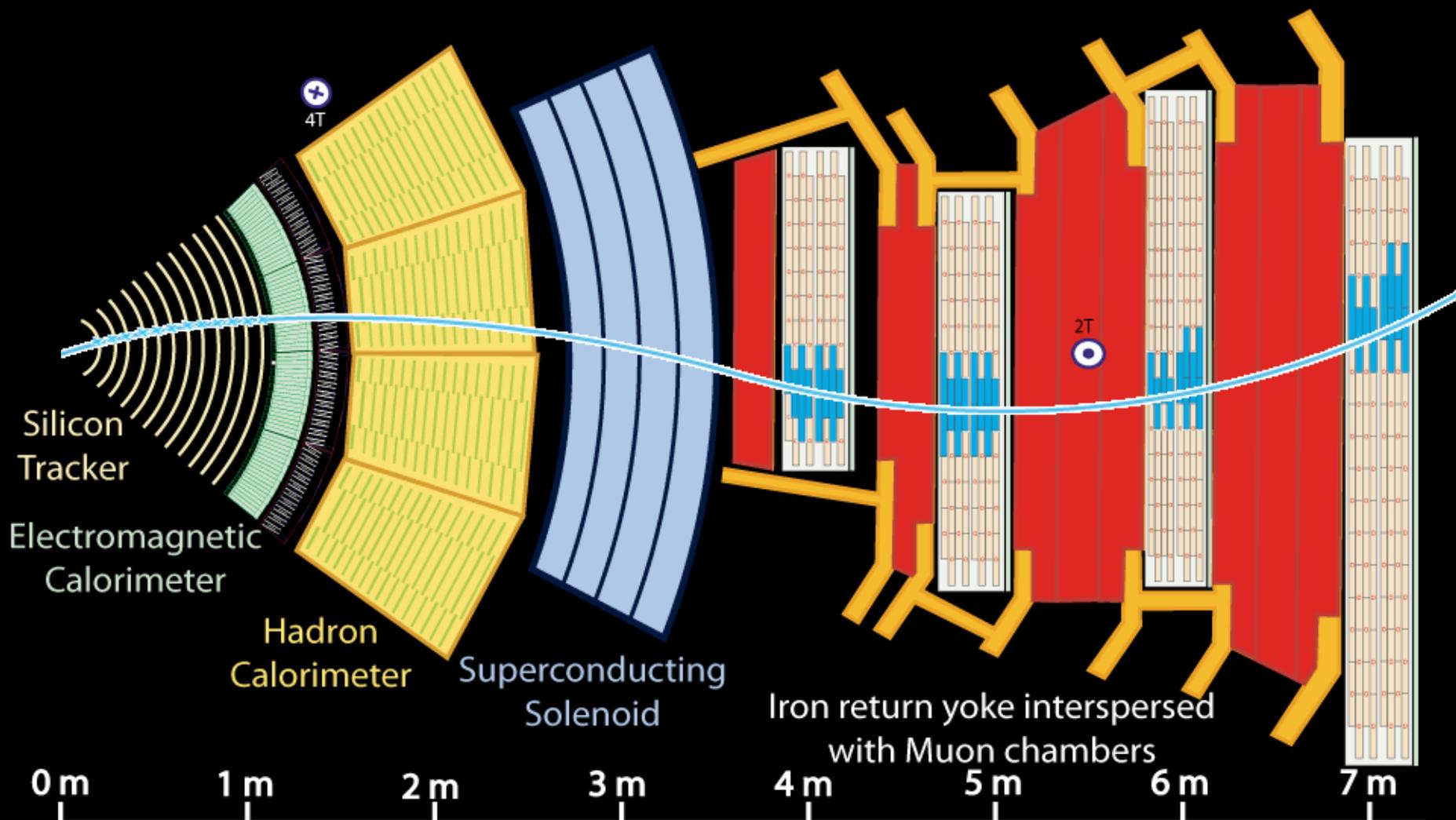
— Muon

— Electron

— Charged Hadron (e.g. Pion)

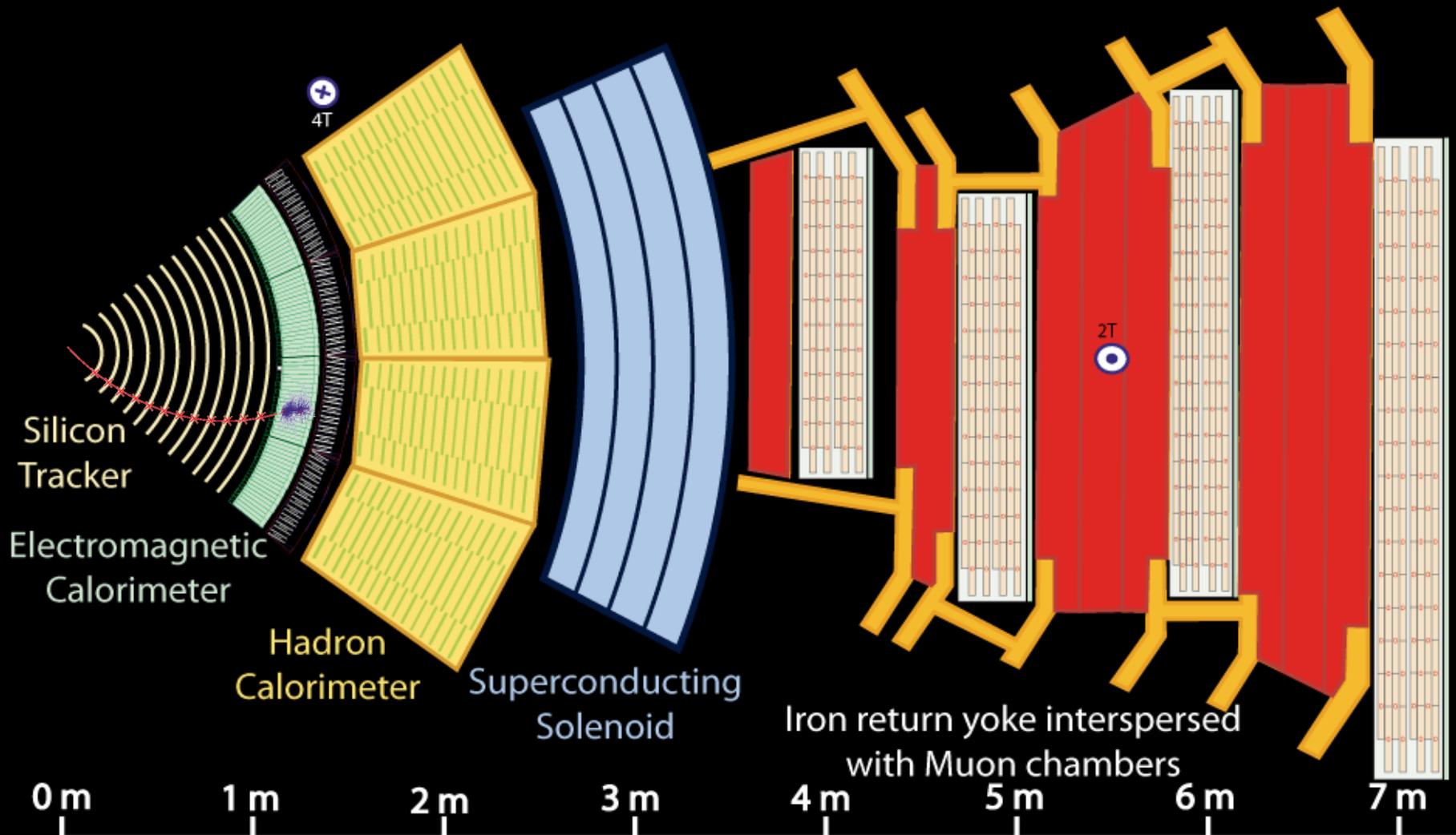
- - - Neutral Hadron (e.g. Neutron)

- - - Photon



Key:

- Muon
- Electron
- Charged Hadron (e.g. Pion)
- - - Neutral Hadron (e.g. Neutron)
- - - Photon



Key:

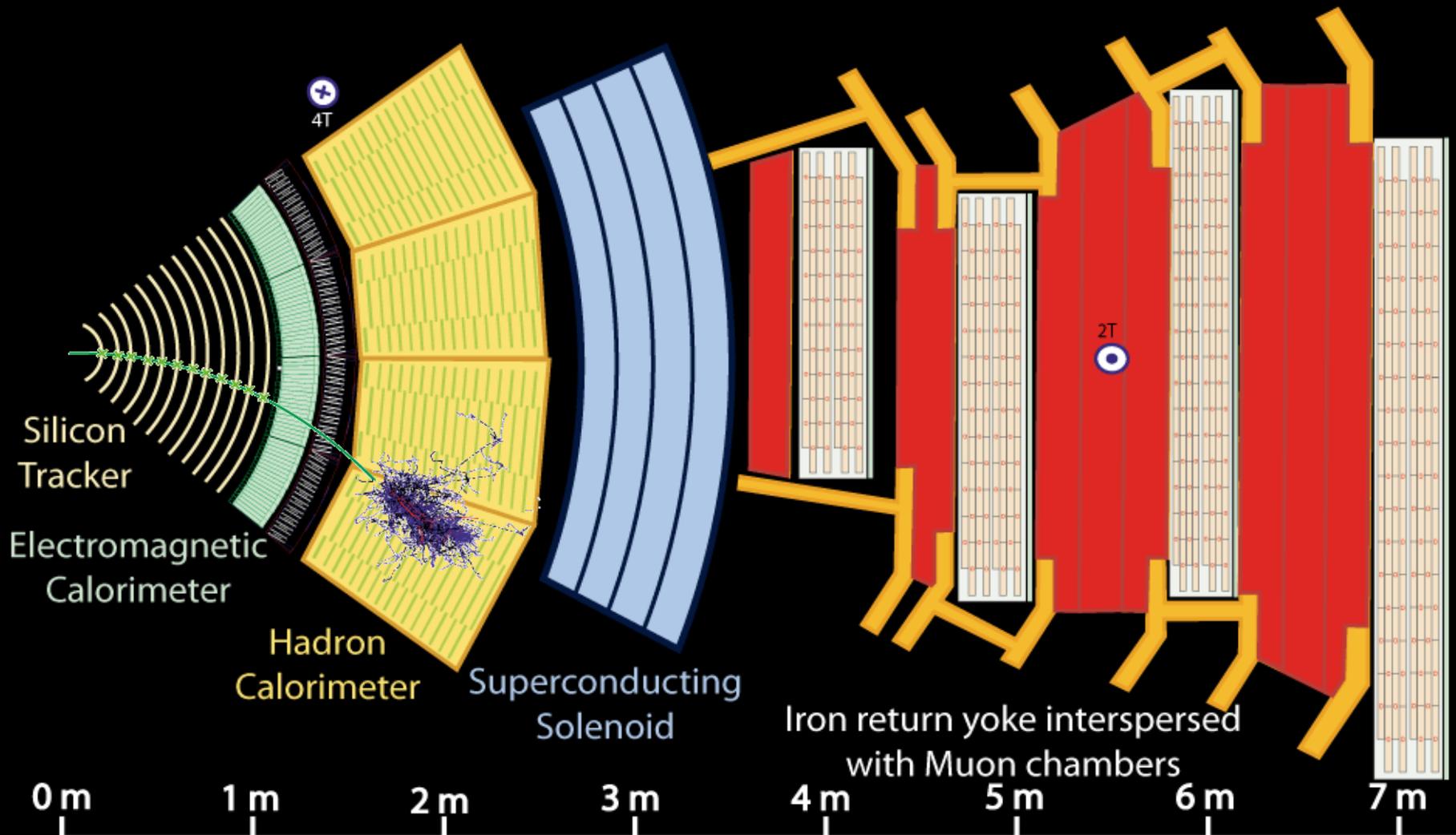
— Muon

— Electron

— Charged Hadron (e.g. Pion)

- - - Neutral Hadron (e.g. Neutron)

- - - Photon



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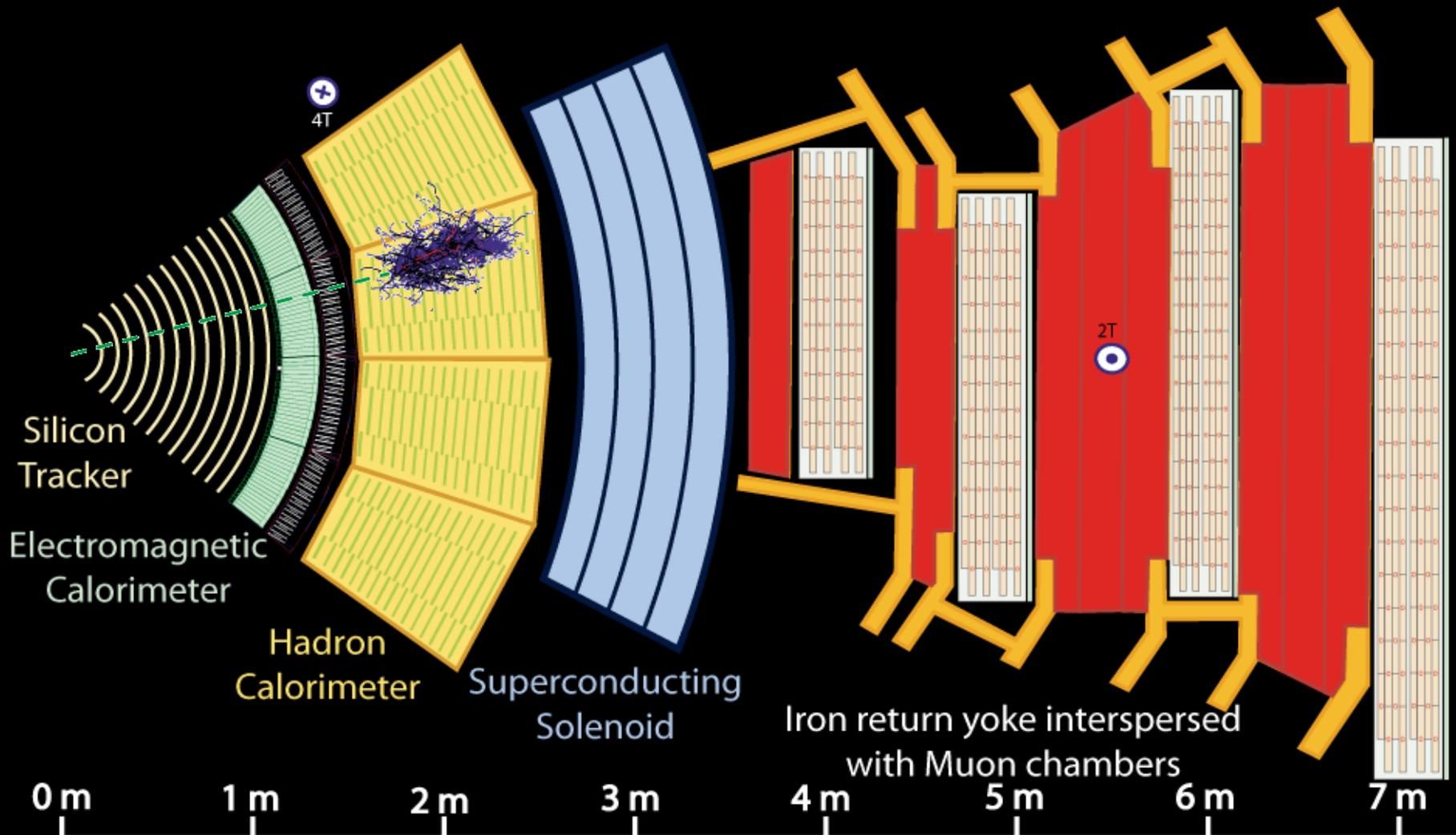
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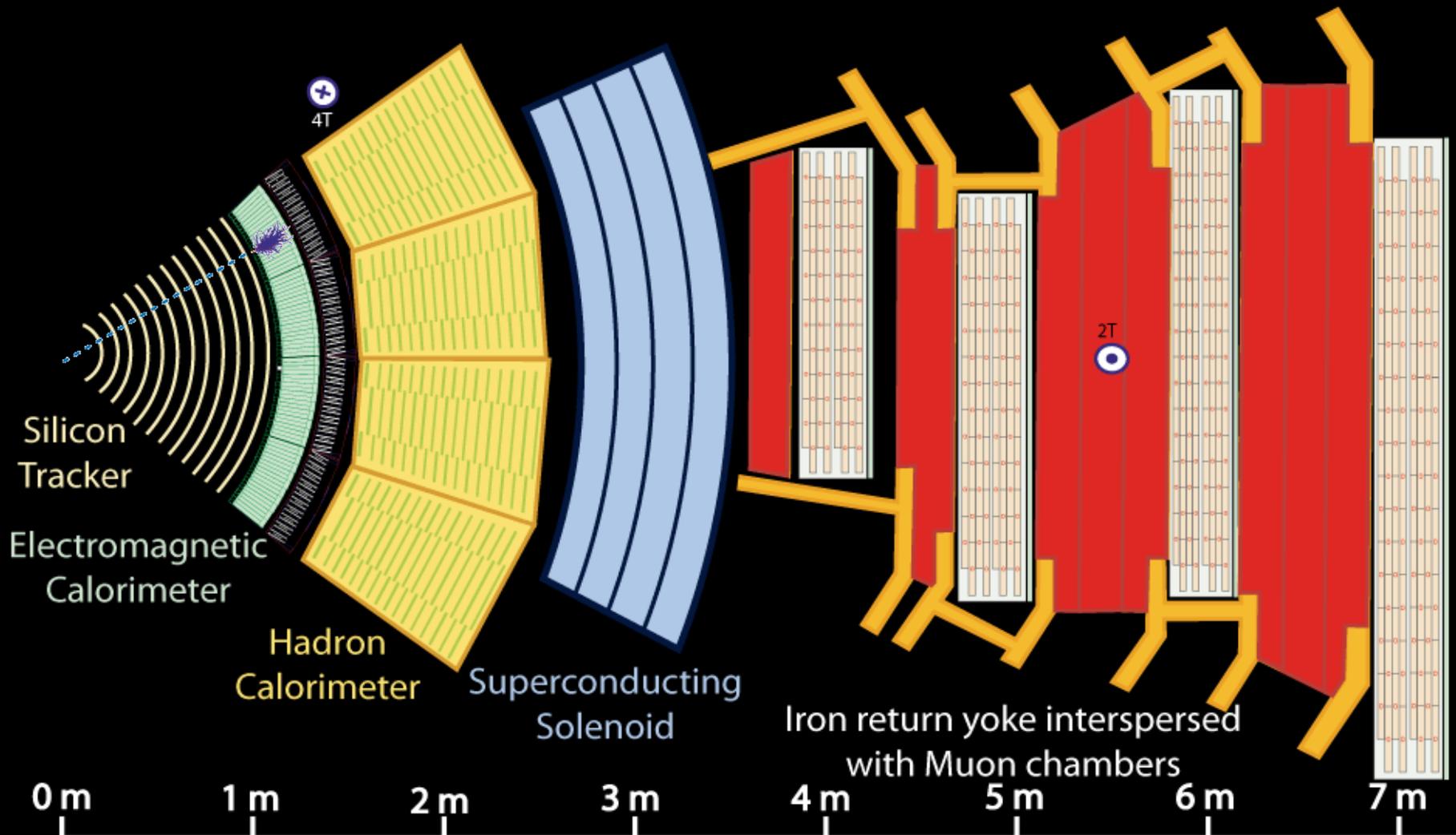
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Key:

— Muon

— Electron

— Charged Hadron (e.g. Pion)

- - - Neutral Hadron (e.g. Neutron)

- - - Photon

CMS assembly

2005



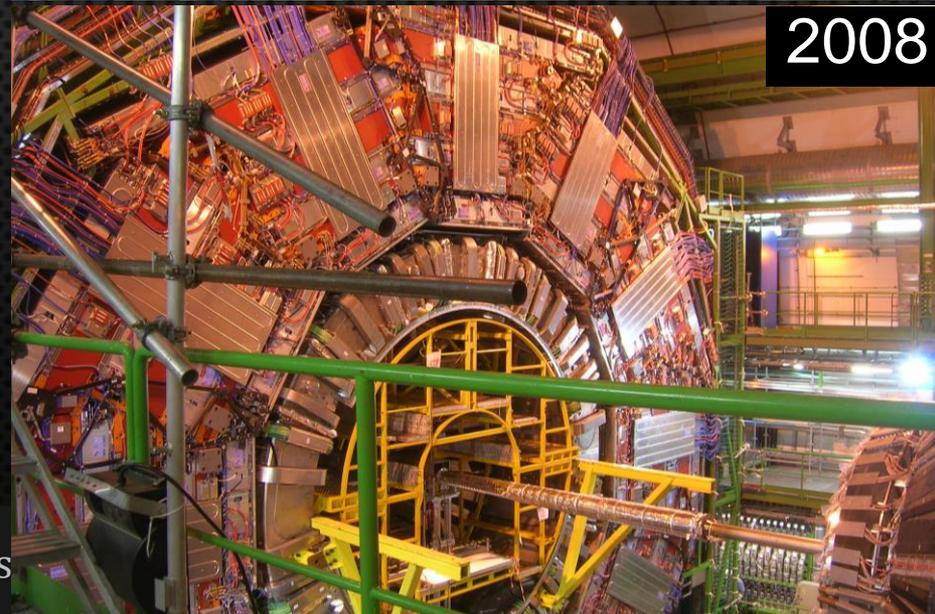
2006



2007

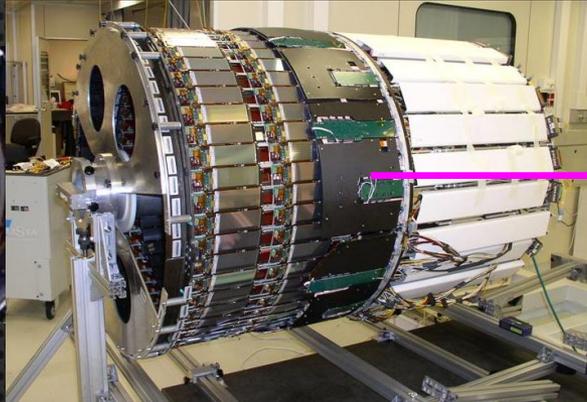
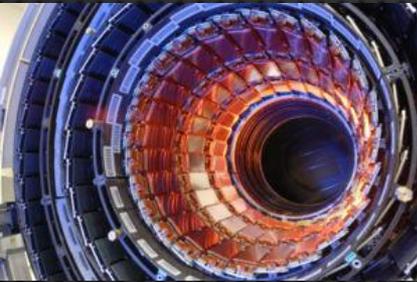


2008

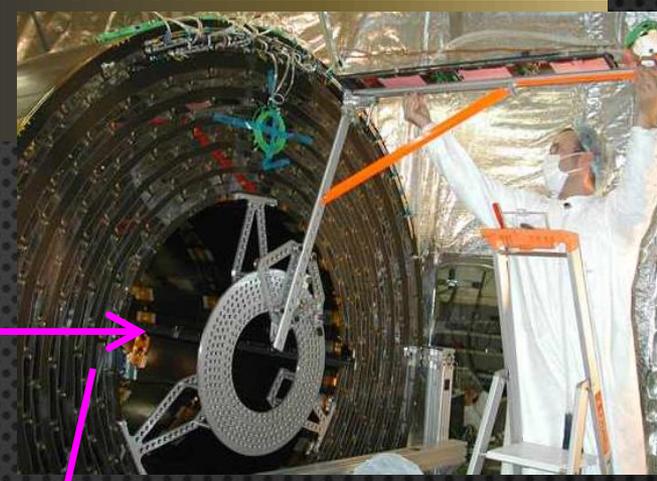


is

CMS Tracker



Goes
inside



CMS tracker uses
2300 square feet of
silicon detectors.

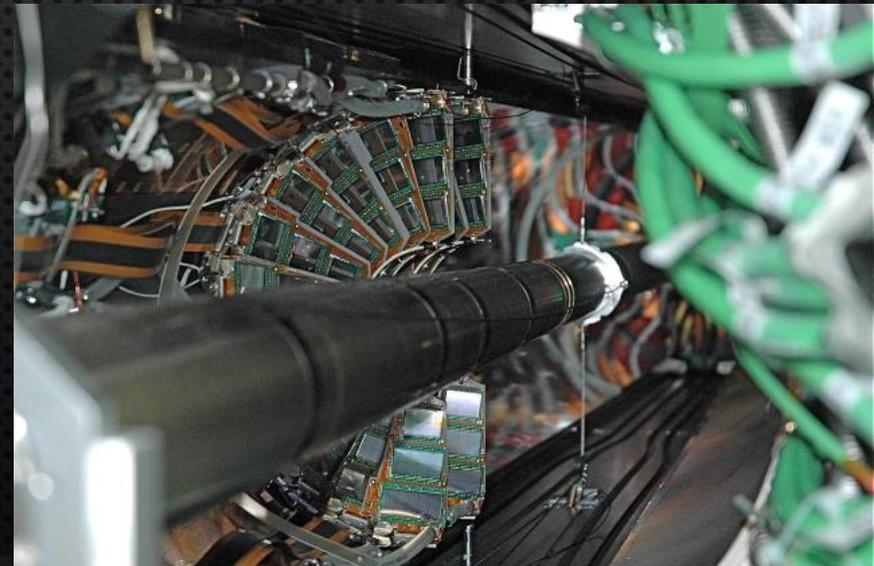
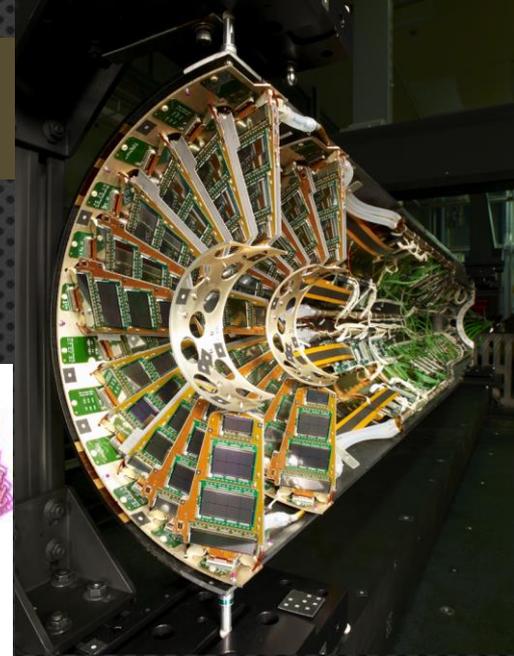
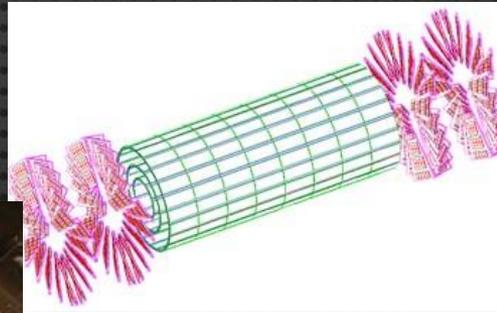
CMS tracker being
inserted into CMS



CMS pixel detector

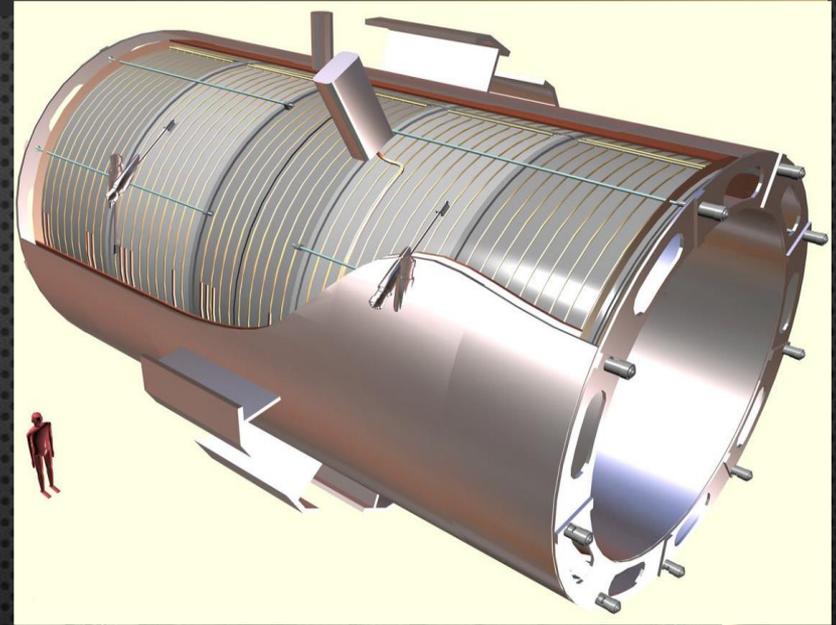
Smallest detector but the most channels. There are 66 million pixels, each $100\ \mu\text{m}$ by $150\ \mu\text{m}$.

Inserting the detector



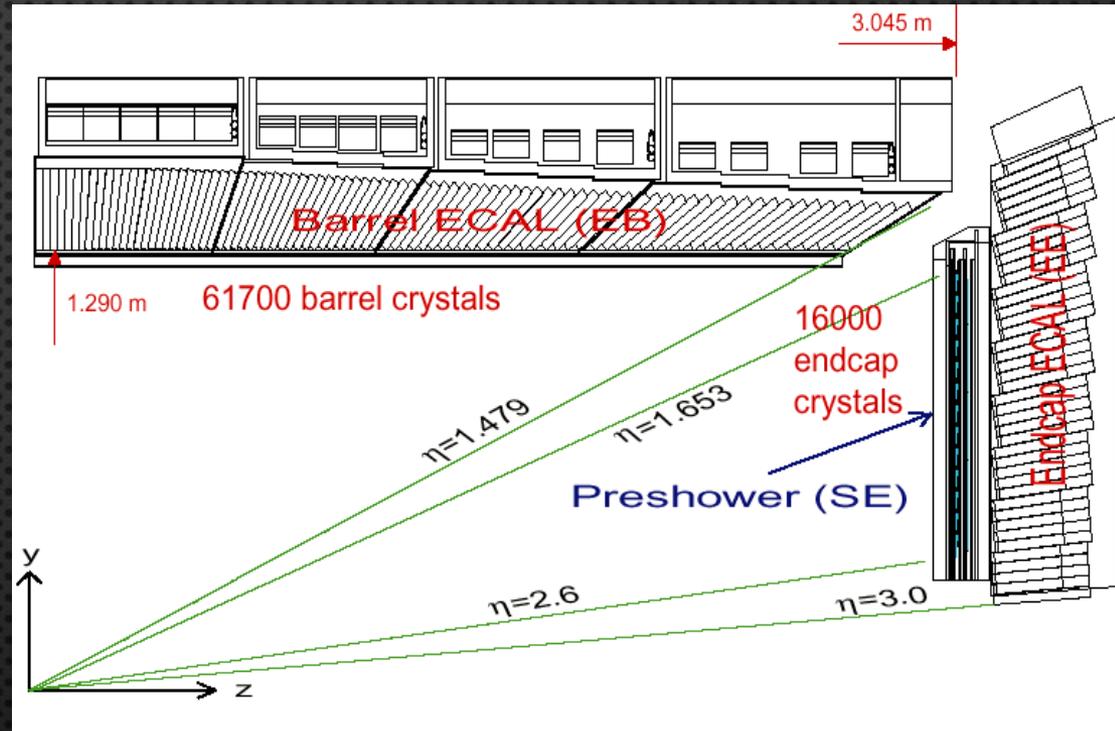
CMS Solenoid

- 3.8 tesla magnet at 4 K
- 6 m diameter and 12.5 m long (largest ever built)
- 220 t (including 6 t of NbTi)
- Stores 2.7 GJ — equivalent to 1300 lbs of TNT
- If magnet gets above superconducting temperature, energy is released as heat – need to plan for the worst
- Bends charged particles allowing tracker to measure momentum



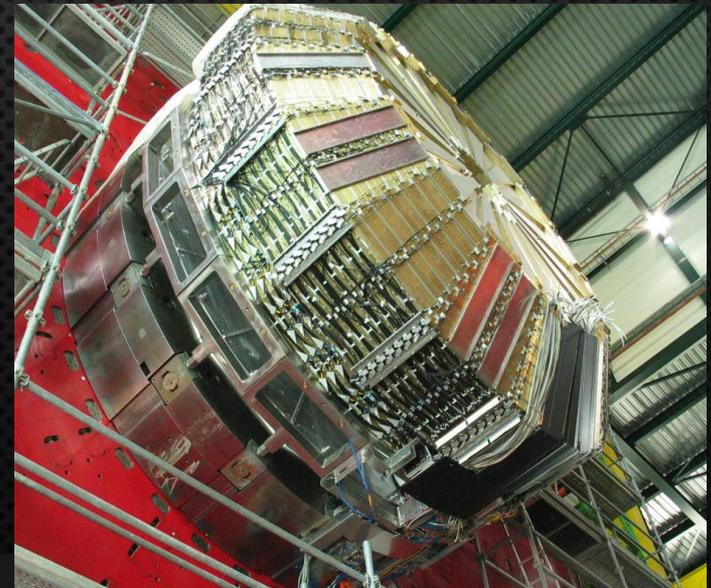
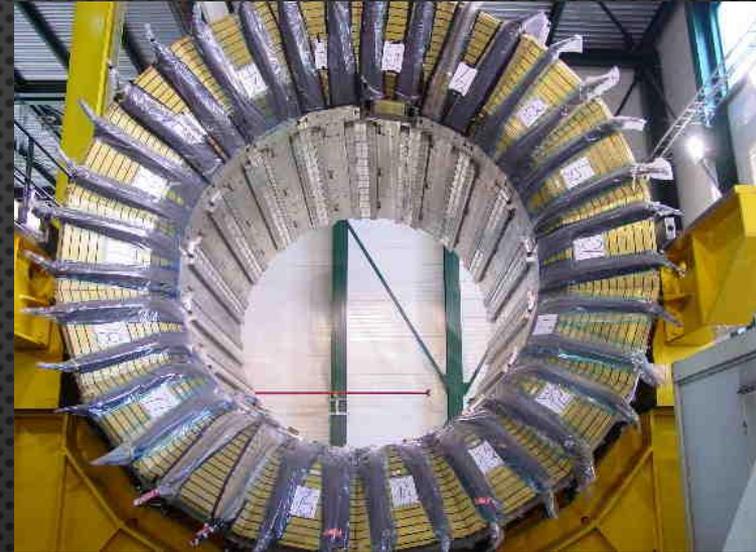
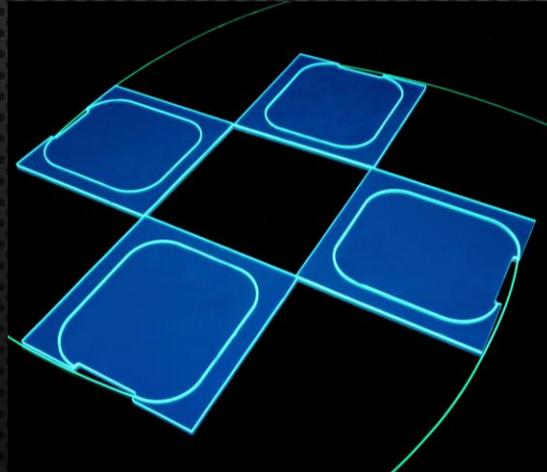
CMS Electromagnetic Calorimeter (ECAL)

- Photons and electrons shower in high Z material
- Homogenous calorimeter
- Lead tungstate (PbWO_4) crystals: $2.3 \times 2.3 \times 23 \text{ cm}^3$
- Radiation hard, dense, and fast
- Low light yield & temperature sensitivity make it difficult
- Magnetic field and radiation require novel electronics APD and VPT



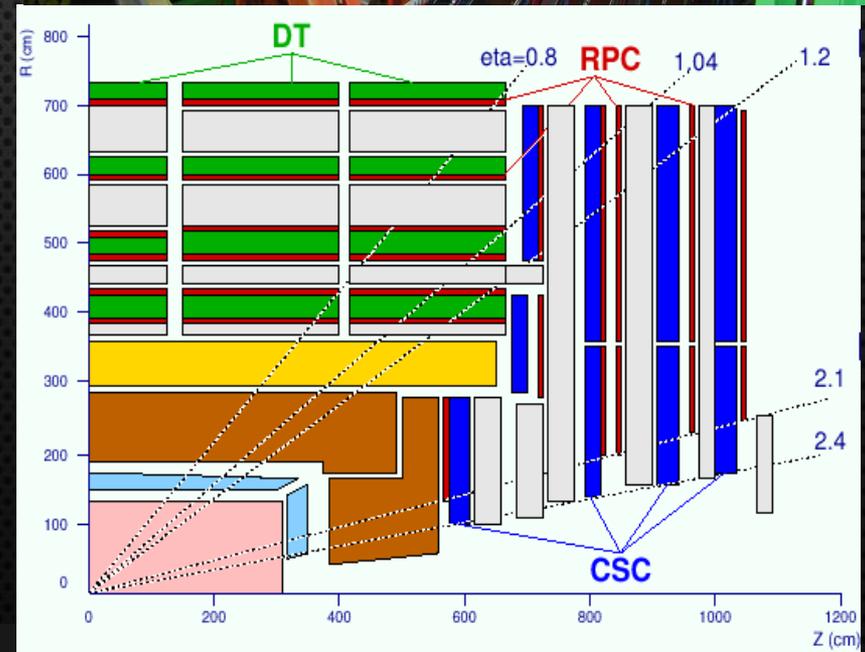
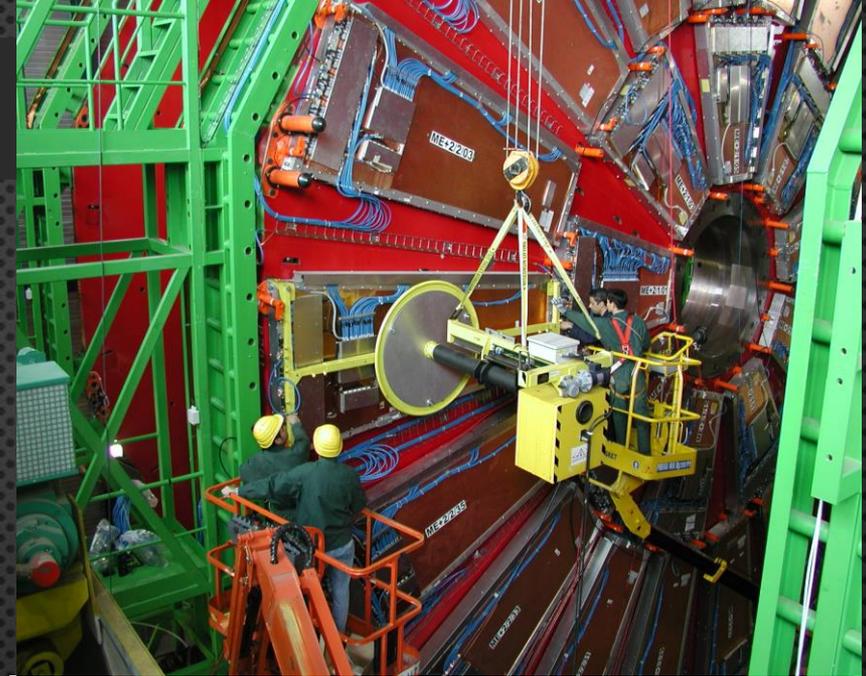
CMS Hadronic Calorimeter (HCAL)

- Sampling calorimeter
- Brass absorber from Russian artillery shells (non-magnetic)
- Scintillating tiles with wavelength shifting (WLS) fiber
- WLS fiber is fed into a hybrid photo-diode (HPD) for light yield measurement



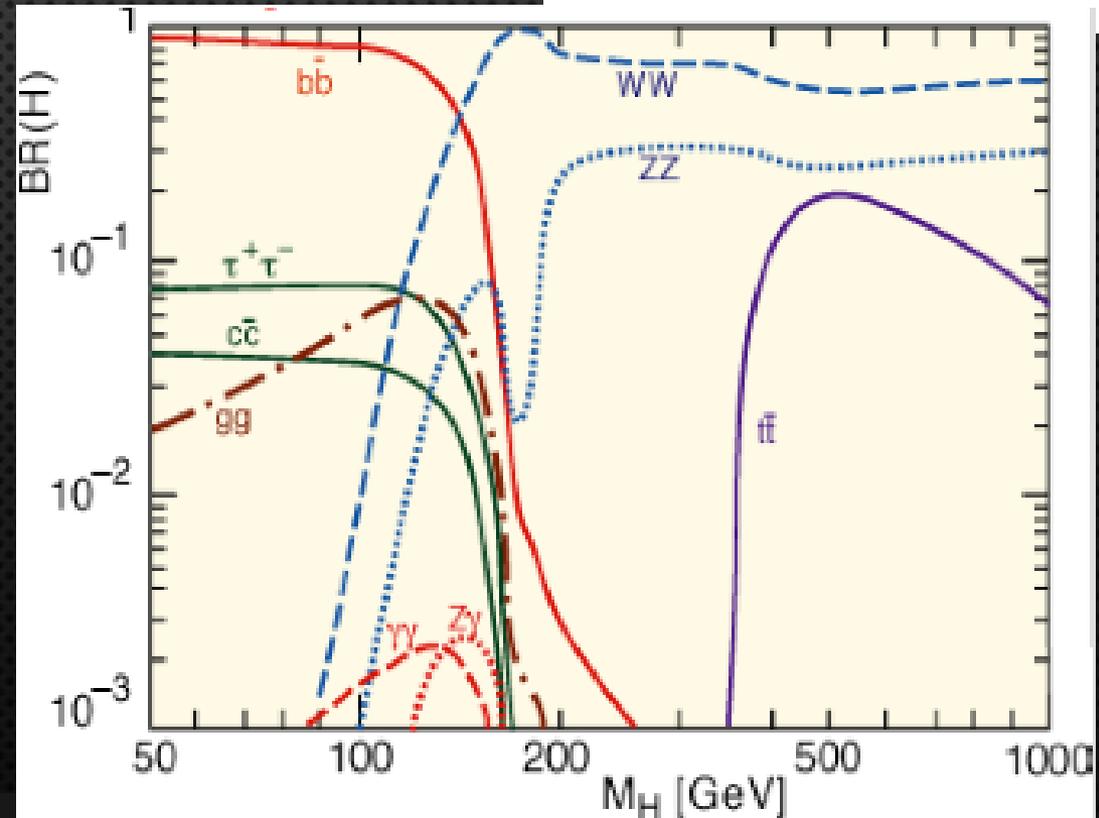
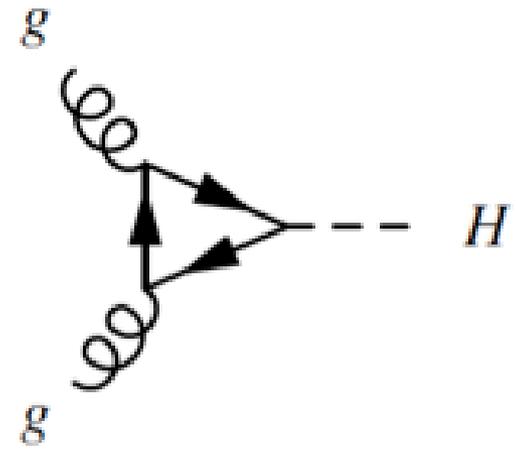
CMS Muon detector

- Muons interact less than other charged particles
- Place detectors after lots of material so that what comes through must be a muon
- Add magnetic field & tracking to find momentum and link with main tracker
- 12000 t of iron is absorber and solenoid flux return
- Three tracking technologies: Drift Tube, Resistive Plate Chamber, & Cathode Strip Chamber



Higgs production and decay

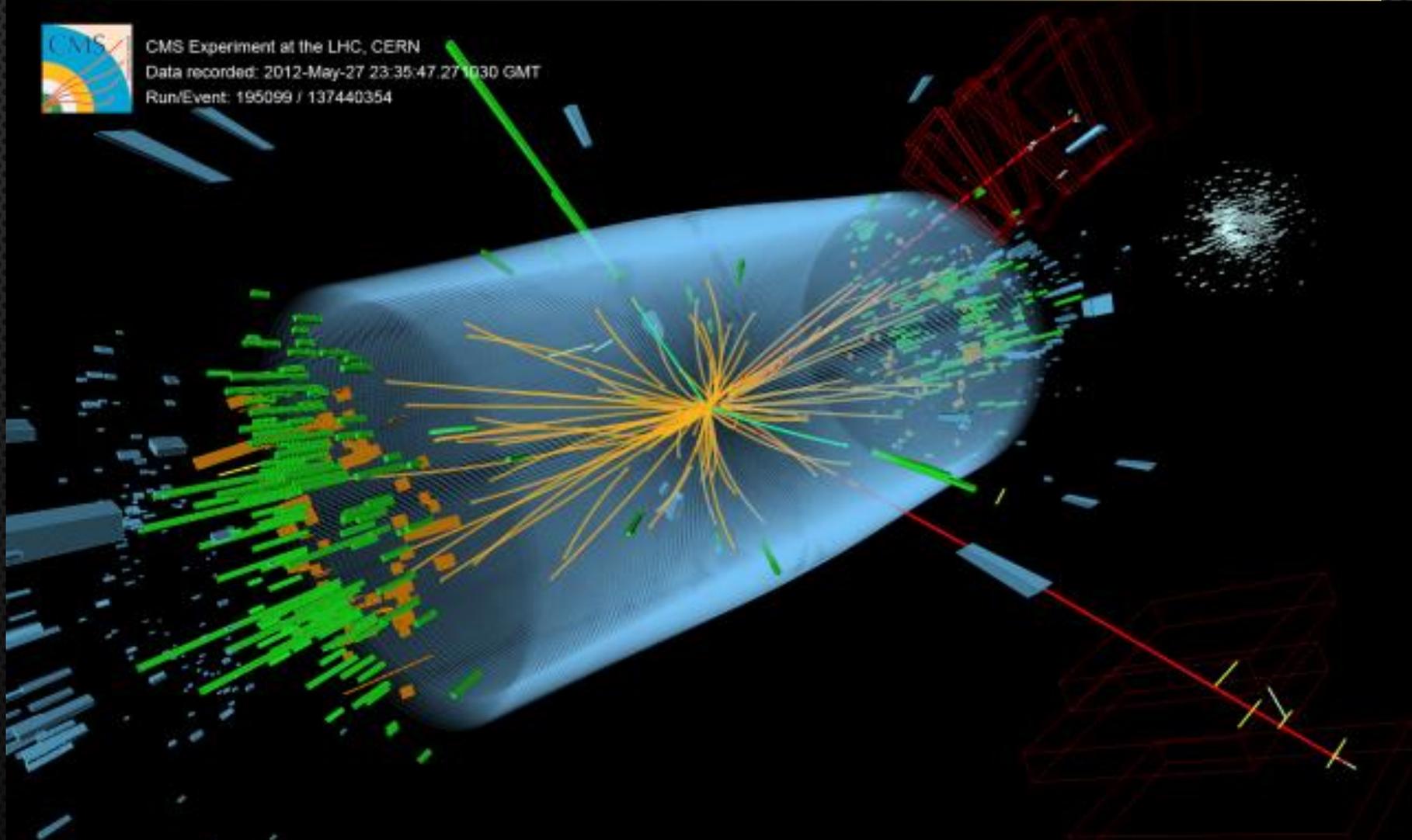
- Two gluons inside the colliding protons fuse to make a Higgs boson.
- The Higgs is unstable so it decays into other particles, some of which also decay.
 - $H \rightarrow \gamma\gamma$ and $H \rightarrow ZZ$ are the most important decay modes.



Possible Higgs decay: $H \rightarrow ZZ \rightarrow ee\mu\mu$



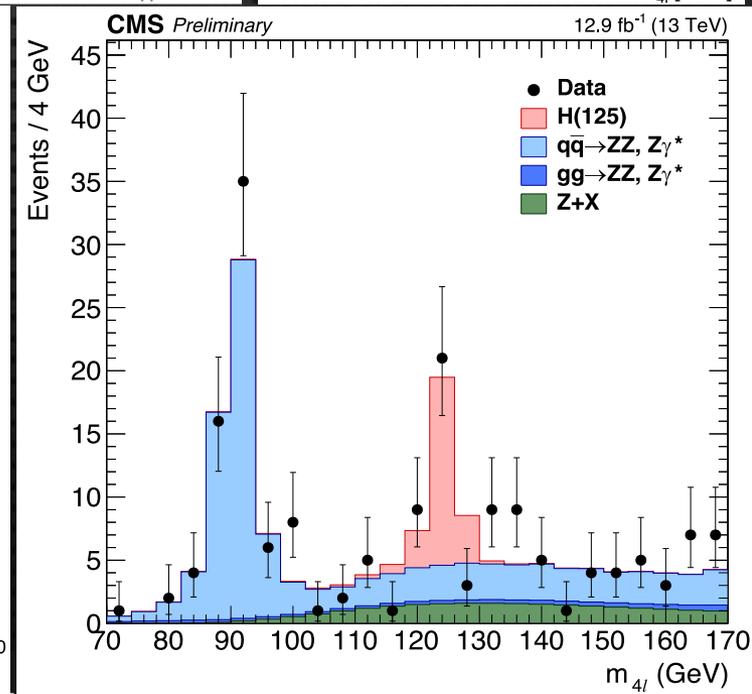
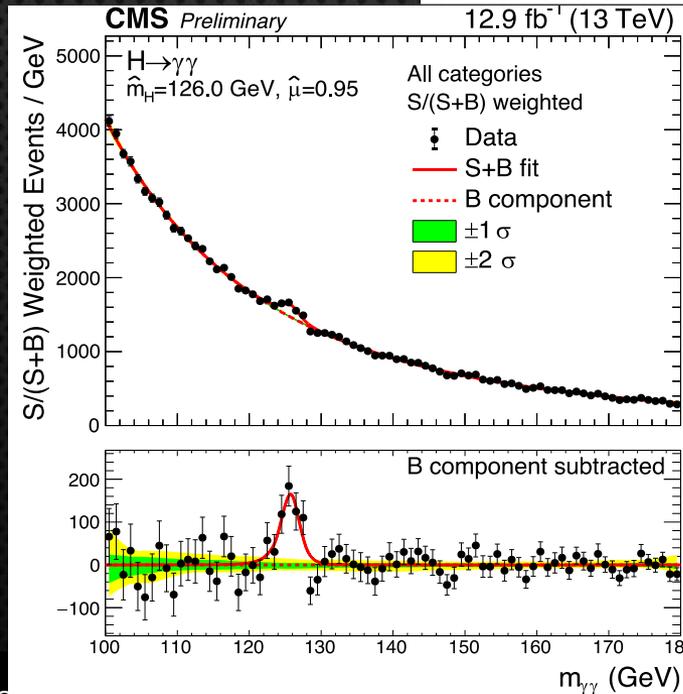
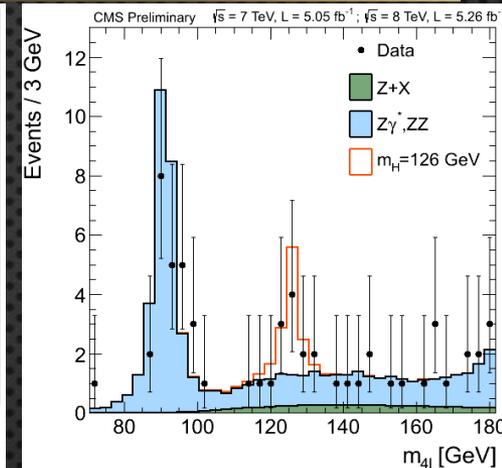
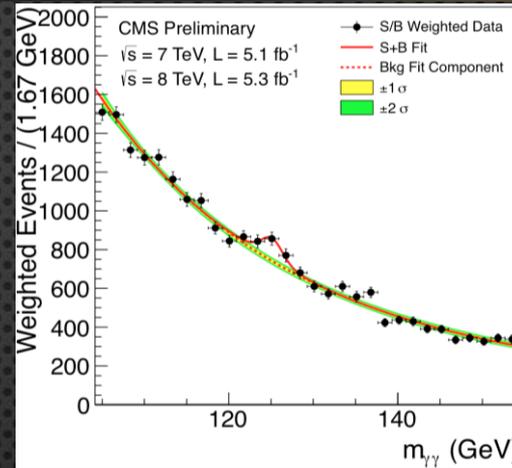
CMS Experiment at the LHC, CERN
Data recorded: 2012-May-27 23:35:47.271030 GMT
Run/Event: 195099 / 137440354



The Higgs discovery

From the July 4, 2012 announcement, evidence of $H \rightarrow \gamma\gamma$ and $H \rightarrow \text{IIIL}$.

Recent results from 2016 data shown in August.

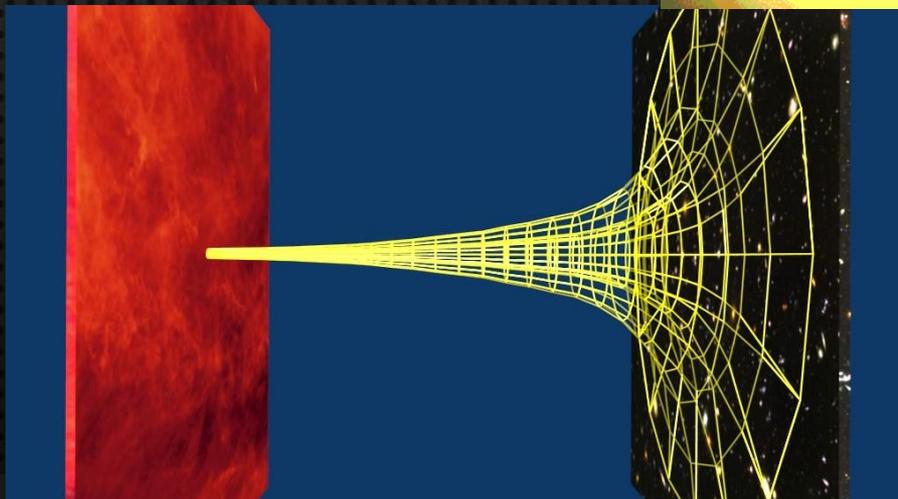
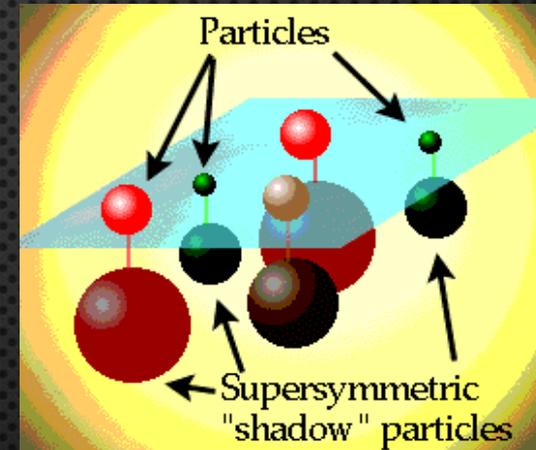


But still lots of questions...

- How come the Higgs boson mass is so light?
- What is the source of dark matter?
- Is there a Grand Unified Theory?
- What is the nature of dark energy?
- What happened to all the antimatter in the universe?
- Why are there 3 generations of quarks and leptons?
- How do neutrinos get their mass?
- Are neutrinos Dirac or Majorana particles?
- What is the quantum theory of gravity?

What are the new theories?

- It is clear that the Standard Model is incomplete since it can't answer all the questions we have. Some of the theories we are testing at the LHC are:
 - Supersymmetry (SUSY)
 - Warped extra dimensions
 - Extra particles (matter or gauge particles)
 - Low(ish) energy string theory
 - Microscopic black holes



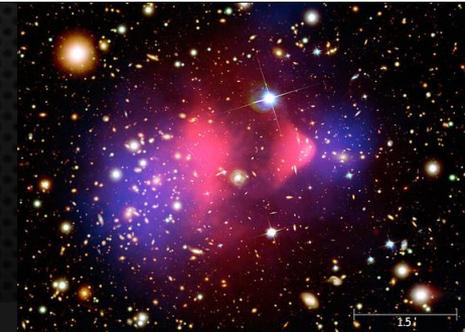
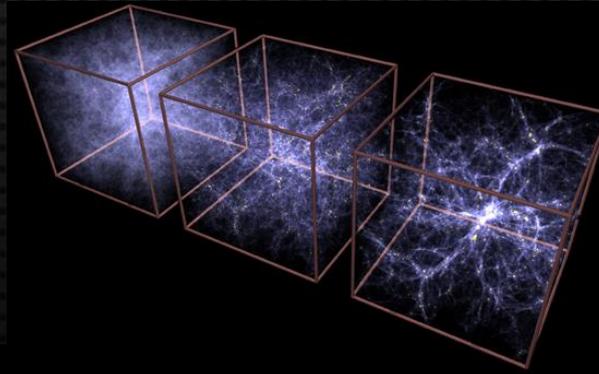
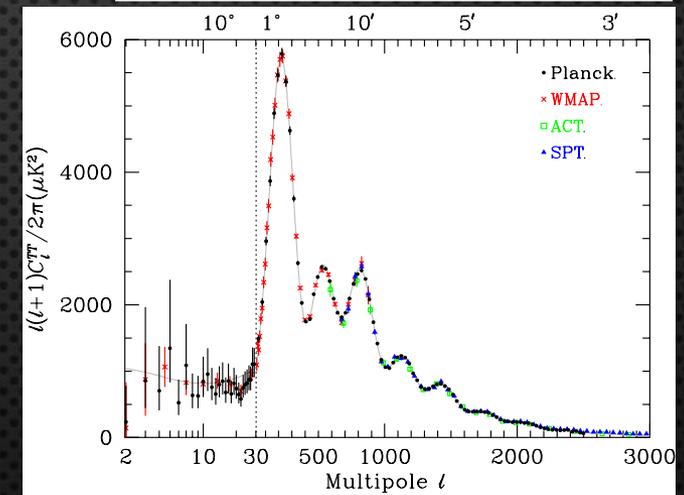
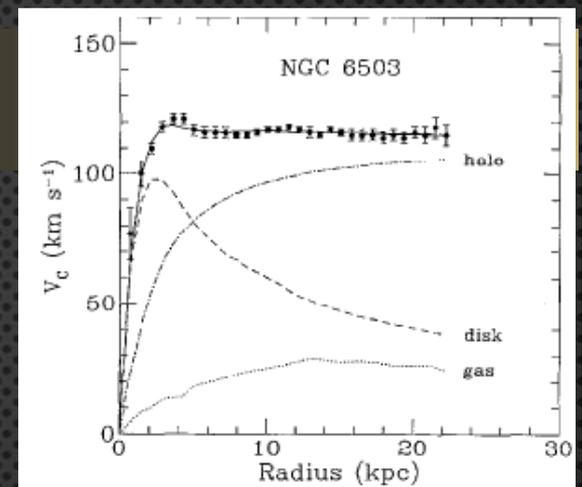
Why is the Higgs boson so light?

- The SM is known to work up to an energy of ~ 1 TeV.
- It is expected to break down before the Planck scale (10^{16} TeV) where quantum gravity should appear.
- So the SM cutoff scale, labeled Λ , should be $< 10^{16}$ TeV
- Calculations of the Higgs mass in the SM contain quantum corrections that should result in a mass near Λ . So it is strange to find it at 0.1 TeV.
- Solutions to the hierarchy problem:
 - The SM is only valid to ~ 1 TeV and there is a new theory above ~ 1 TeV.
 - There is a fine tuning of the quantum corrections at the level of 1 part in 10^{16} (actually 1 part in 10^{32} since they are quadratic).



Evidence for dark matter

- Galactic rotation curves indicate invisible matter in a spherical halo.
- Cosmic microwave background finds more than normal matter.
- Gravitational lensing of Bullet Cluster shows dark matter behaves differently than ordinary matter.
- Large scale structure of the universe only makes sense with dark matter.

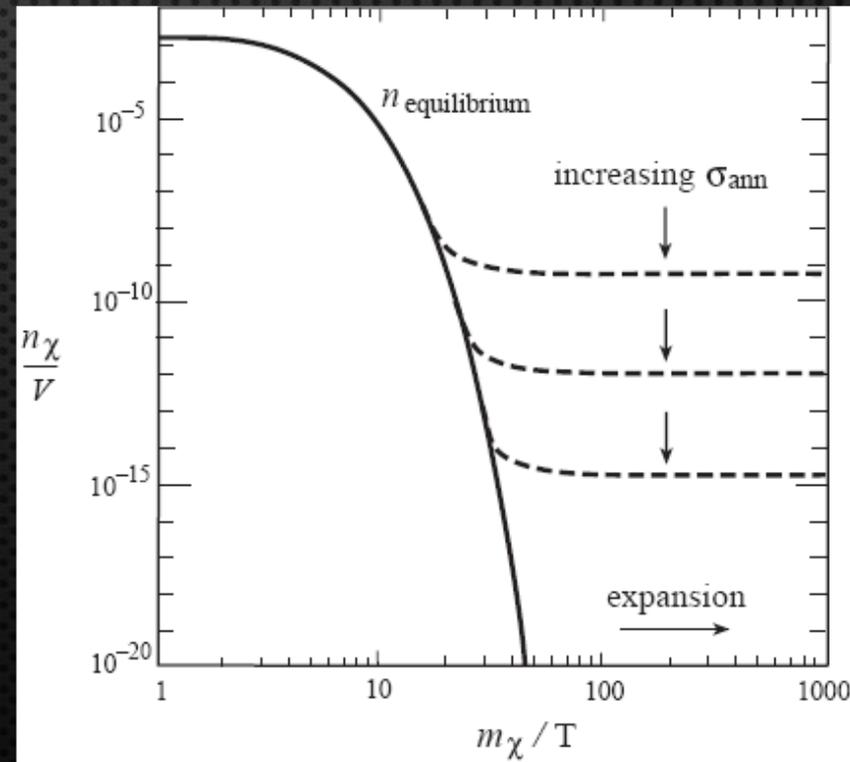


Some dark matter possibilities

- **Primordial black holes** – Created during the big bang before protons are created.
- **Axions** – Light particles with very little interaction with normal matter; proposed to explain lack of CP violation in the strong interaction.
- **Sterile neutrinos** – Masses around keV but not coupling to the Z or other neutrinos.
- **Weakly Interacting Massive Particles (WIMPs)**
 - Interact via weak interaction (and gravity) only
 - Should have mass $\sim 0.1\text{--}1$ TeV

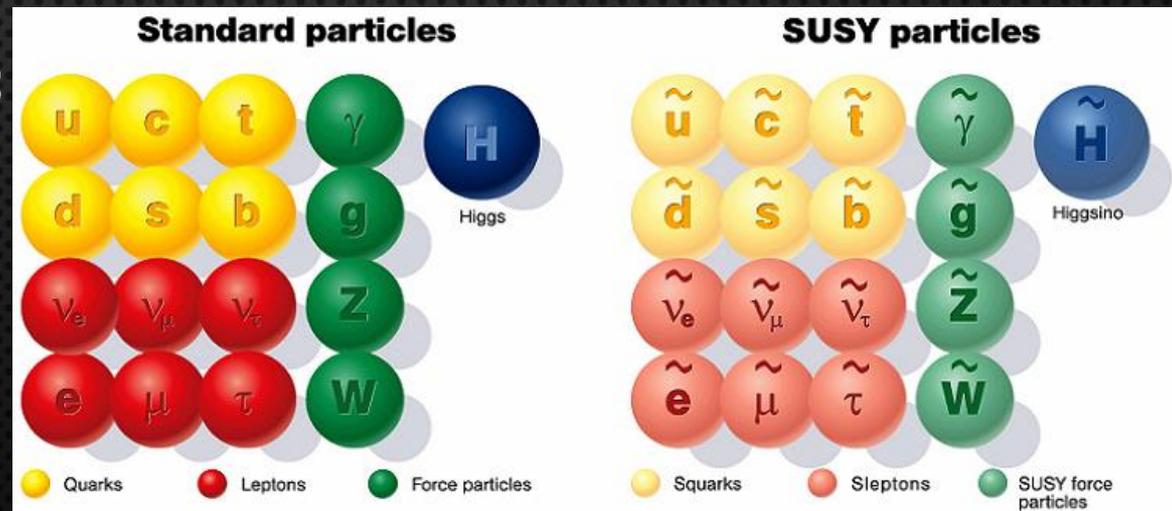
WIMPs seem to be an ideal solution

- During Big Bang, universe is cooling and expanding.
- When $T > M_\chi$, the number density of χ particles is constant as creation and annihilation is in equilibrium.
- When $T < M_\chi$, density drops off exponentially (Boltzman).
- As universe expands, particles can't find each other to annihilate: "freeze-out".
- The cross section needed to get the correct amount of dark matter is the same as the weak interaction. Called the "WIMP miracle".



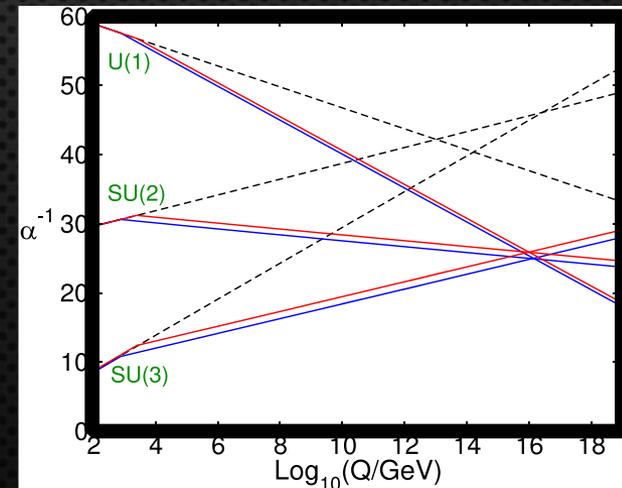
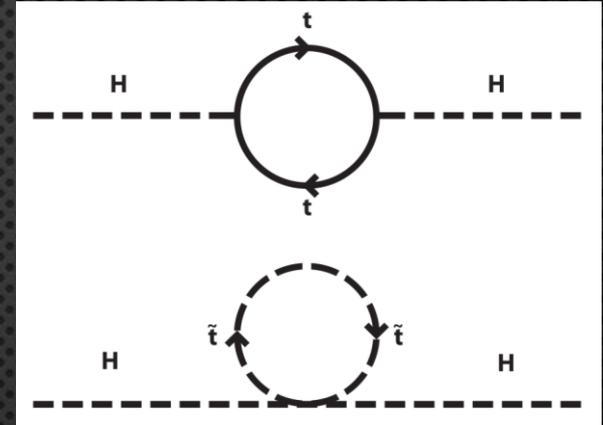
Enter Supersymmetry (SUSY)

- Proposes an additional symmetry in nature between bosons and fermions.
- Each SM particle has a partner particle (sparticle) with same quantum numbers except spin differs by $\frac{1}{2}$.
- Cool sparticle names like sbottom, selectron, wino.
- Consider R -parity conserving models today: sparticles are produced in pairs and decay to the lightest supersymmetric particle (LSP).



Three good features of SUSY

1. Can solve hierarchy problem: Quantum corrections from SM particles canceled by the sparticles.
2. Can provide a dark matter candidate. A neutral LSP will be a WIMP.
3. The strong, weak, and E&M forces have different strengths (couplings) that change (run) with energy. When the effects of SUSY are included, these couplings converge at 10^{16} GeV, keeping alive the hope for a Grand Unified Theory.

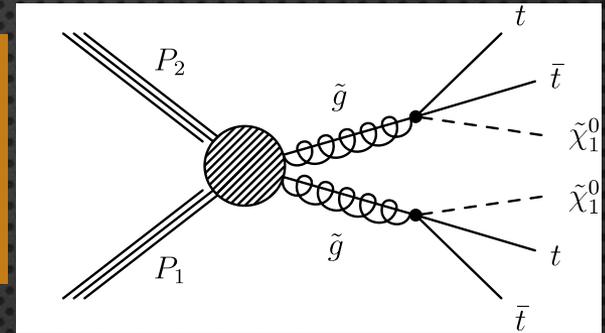


The less great features of SUSY

- SUSY itself is a well defined theory.
- If supersymmetry was a good symmetry, the sparticle masses would be the same as the particle masses.
- Since we haven't found any sparticles, there must be a symmetry breaking mechanism.
- This is no problem (already seen – electroweak symmetry breaking is what gives mass to most SM particles).
- But we have no idea how the symmetry breaking is done so this opens up all sorts of possibilities – not very predictive.
- Need to look as broadly as possible for signs of SUSY.

Search for SUSY in hadronic modes

Search for SUSY involving production via the strong interaction (larger cross section) so produce gluinos or squarks.



Always have 2 LSPs, which exit the detector undetected. In a hermetic detector, this can be seen as large missing transverse energy (**MHT**).

High mass events produce large amounts of energy seen in the detector (**HT**).

Each top quark decays to b quark (creating a b jet) and W boson (decaying to two jets 67% of time). Results in many jets (**N_{jet}**) and b jets (**N_b**).

SM backgrounds for search

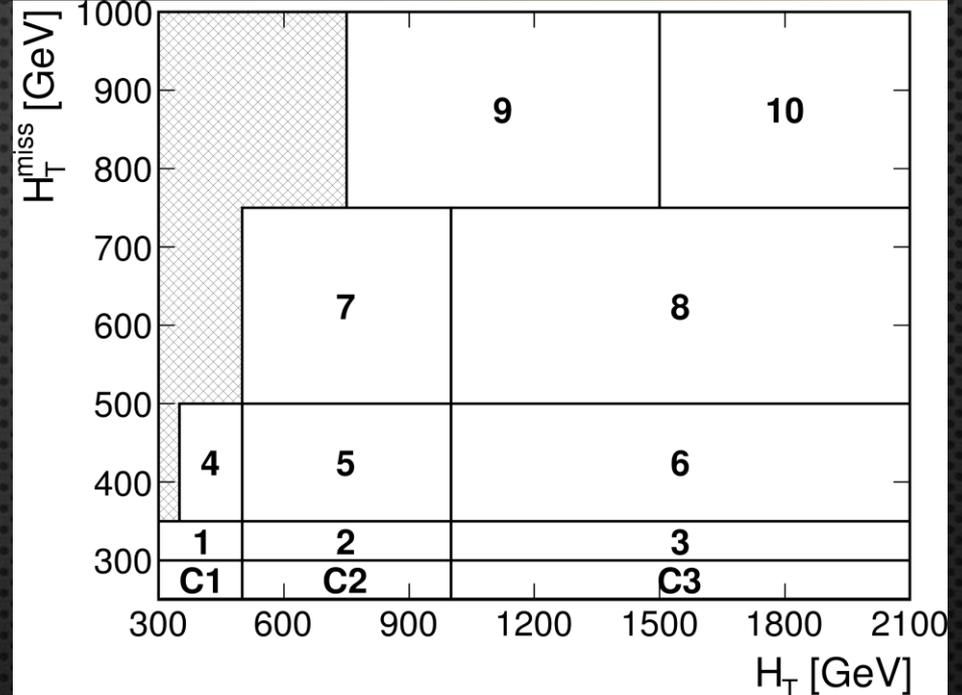
- **QCD** : Direct production of many jets with the energy of jet(s) mismeasured, leading to MHT.
- **Lost-lepton** and **hadronic τ** : Production of W and jets where W decays to $e/\mu/\tau$ and ν . Neutrino creates MHT. Reduced by rejecting events with $e/\mu/\tau$ but some may slip through. W bosons can be produced directly or from top quark decays.
- **$Z \rightarrow \nu\nu$** : Production of Z and jets where Z decays to neutrinos, producing MHT.

These backgrounds are all estimated by using control regions (distinct from signal regions).



Signal region divided into bins

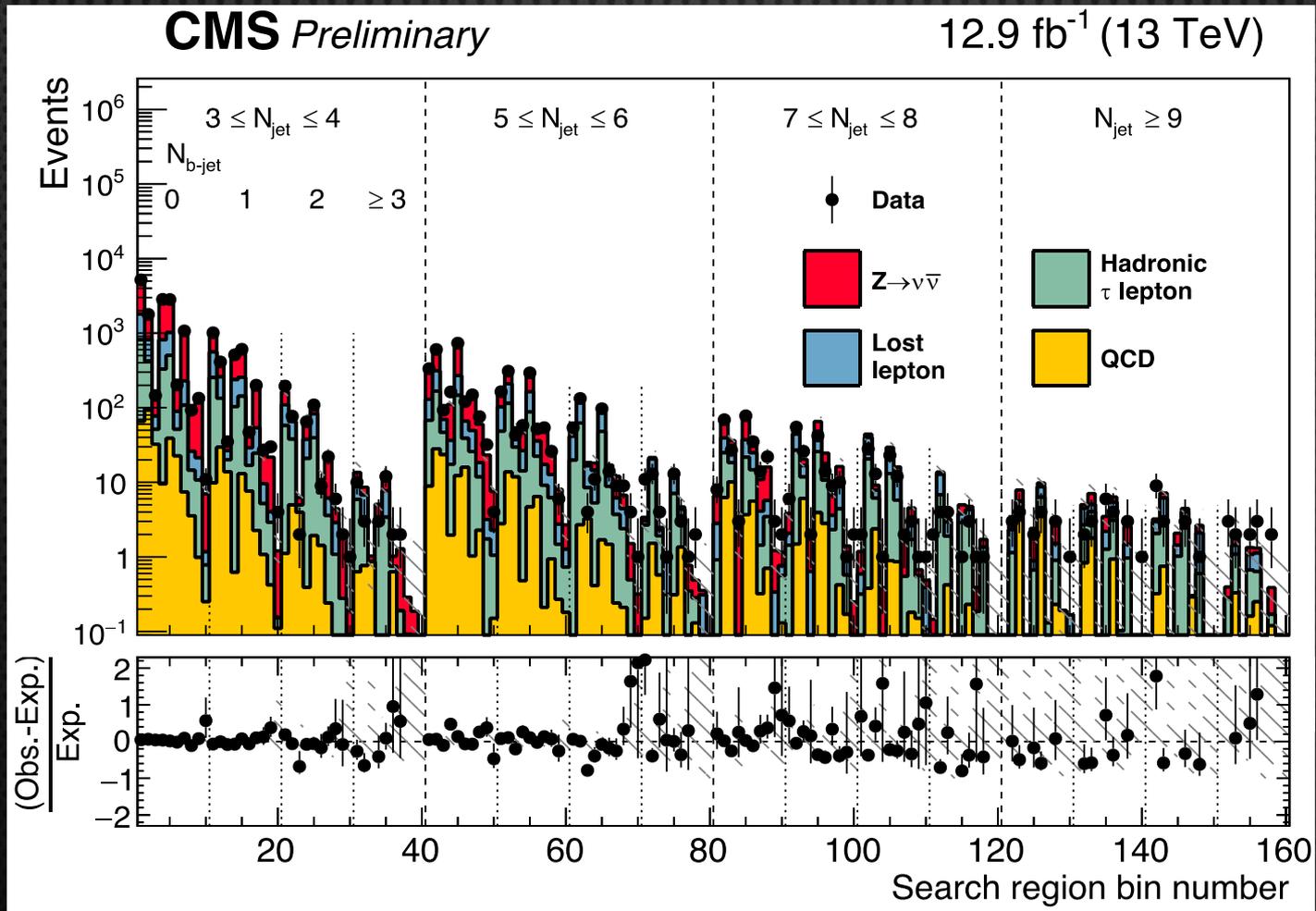
- The HT-MHT plane divided into 10 bins.
- 4 bins of N_{jet} : 3-4, 5-6, 7-8, 9+.
- 4 bins of N_b : 0, 1, 2, 3+.
- Total of 160 bins.



Binning increases the sensitivity and flexibility of the analysis.

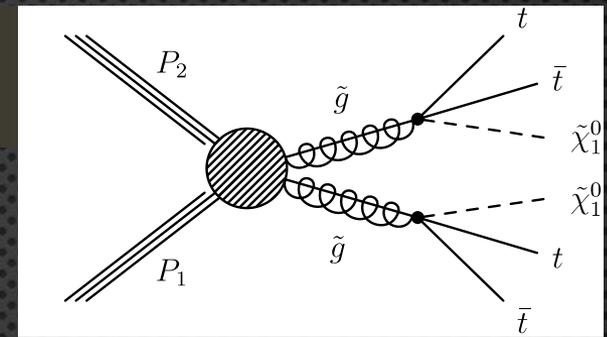
Search results

Data is consistent with estimated backgrounds. No sign of SUSY



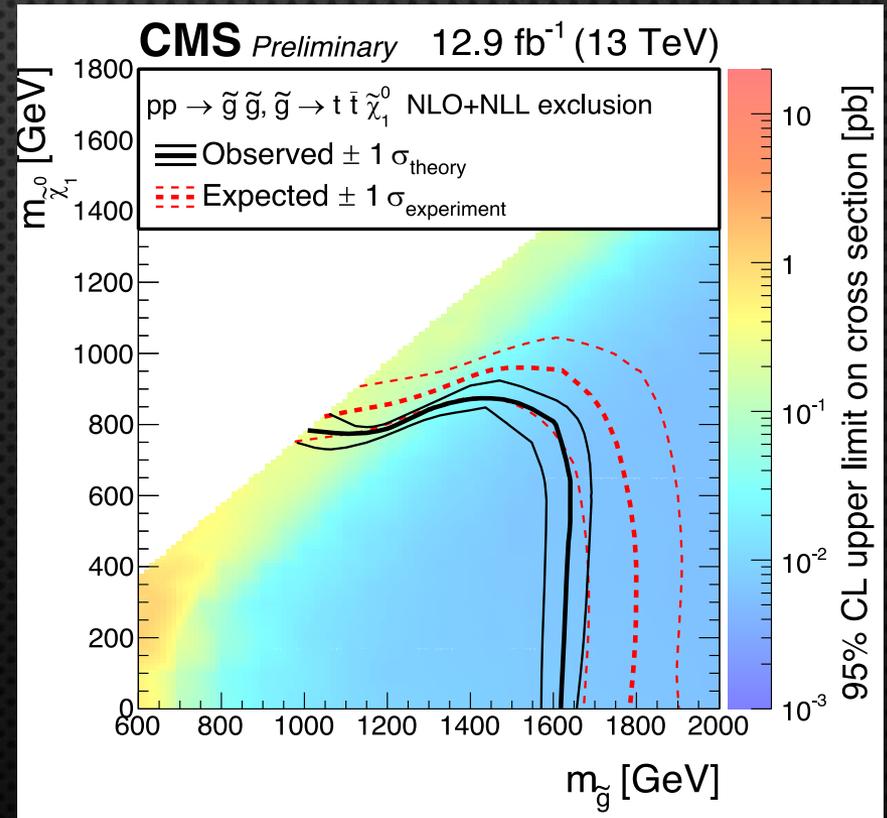
Limits on SUSY production

Results are presented as regions that are excluded.



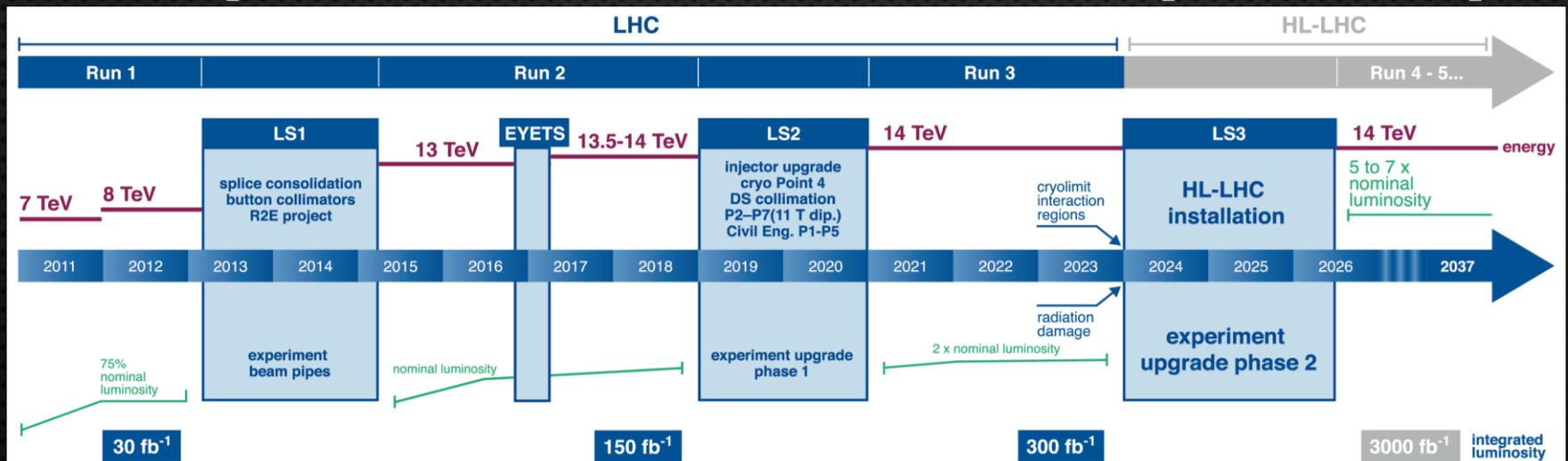
Expected limits are what would be obtained if the observed data exactly matched the prediction.

Observed limits in this case are not as good. Could be background was underestimated, statistical fluctuation, or the first sign of SUSY (probably not).



What does the future hold?

- Analyzing remaining data from 2016 (35 fb⁻¹ total).
- By 2023, this will increase to 300 fb⁻¹. Will greatly expand the reach of our searches for new physics.
- After major upgrades of LHC and detectors, will collect 3000 fb⁻¹ from 2026-2037 for another big increase in sensitivity. Upgraded detectors are being designed now.
- Also expand searches to include more SUSY parameter space.

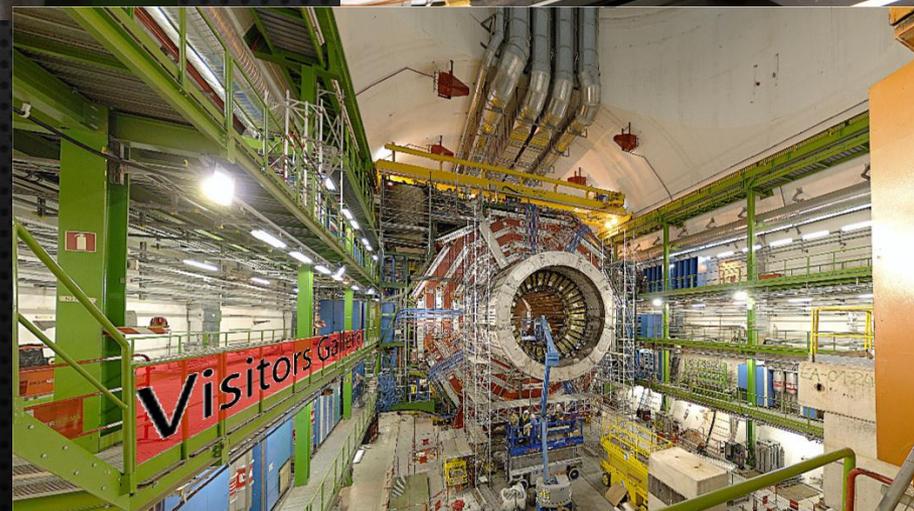
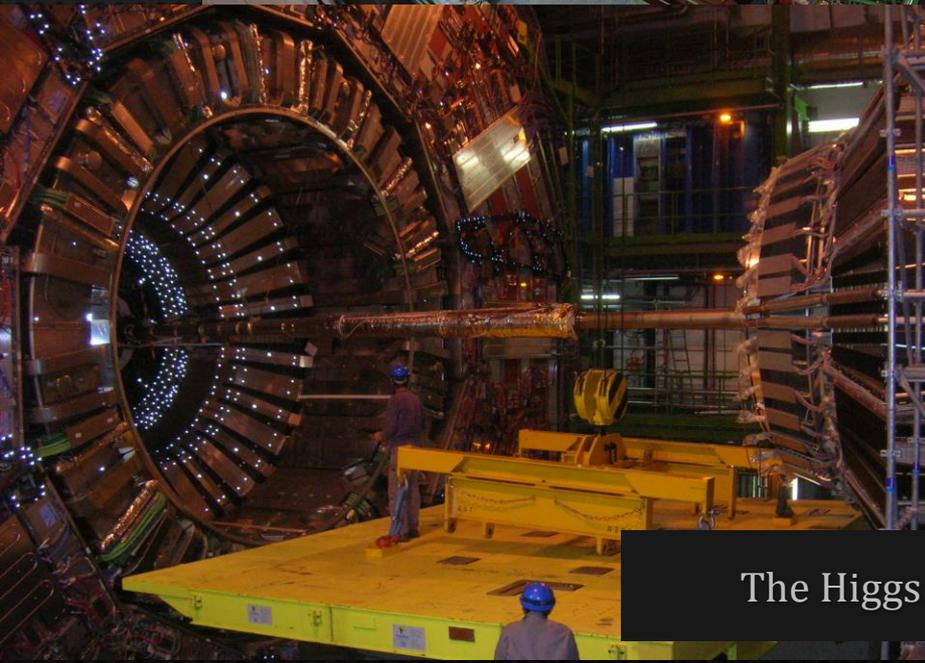
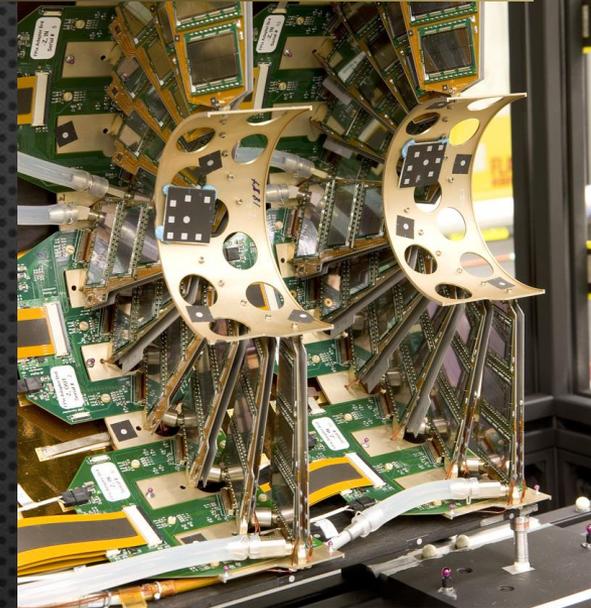


Summary

- The Higgs particle has been discovered, nearly 50 years after its prediction. This completes the Standard Model.
- There are still many open questions in particle physics.
- It is hoped that experiments at the LHC will provide some of the answers.
- So far we haven't found any new physics but during the next 20 years, we should collect 100 times more data that may lead to a revolution in our understanding.

Backup

More photos to end on...



Searching for supersymmetry

A few rules for most SUSY searches:

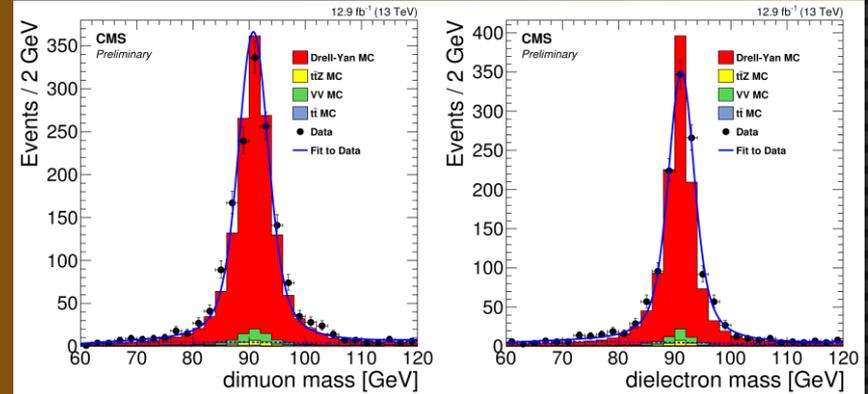
- squarks and gluinos are predominantly produced because they couple to the strong force
- SUSY particles produced in pairs (R-parity conservation)
- Decay cascades end with lightest supersymmetric particle (LSP) which escapes detection.
 - Cascades produce jets & leptons (electrons, muons, taus)
 - The escape of LSP's results in missing energy (MET).

These rules suggest search strategies:

0-leptons	1-lepton	OSDL	SSDL	≥3 leptons	2-photons	γ+lepton
Jets + MET	Single lepton + Jets + MET	Opposite-sign di-lepton + jets + MET	Same-sign di-lepton + jets + MET	Multi-lepton	Di-photon + jet + MET	Photon + lepton + MET

Estimate of $Z \rightarrow \nu\nu$

Reconstruct $Z \rightarrow ee$ and $Z \rightarrow \mu\mu$ events to estimate $Z \rightarrow \nu\nu$. Just need to account for efficiency. Unfortunately, there are more $Z \rightarrow \nu\nu$ events than $Z \rightarrow ee + Z \rightarrow \mu\mu$ events so statistically limited.



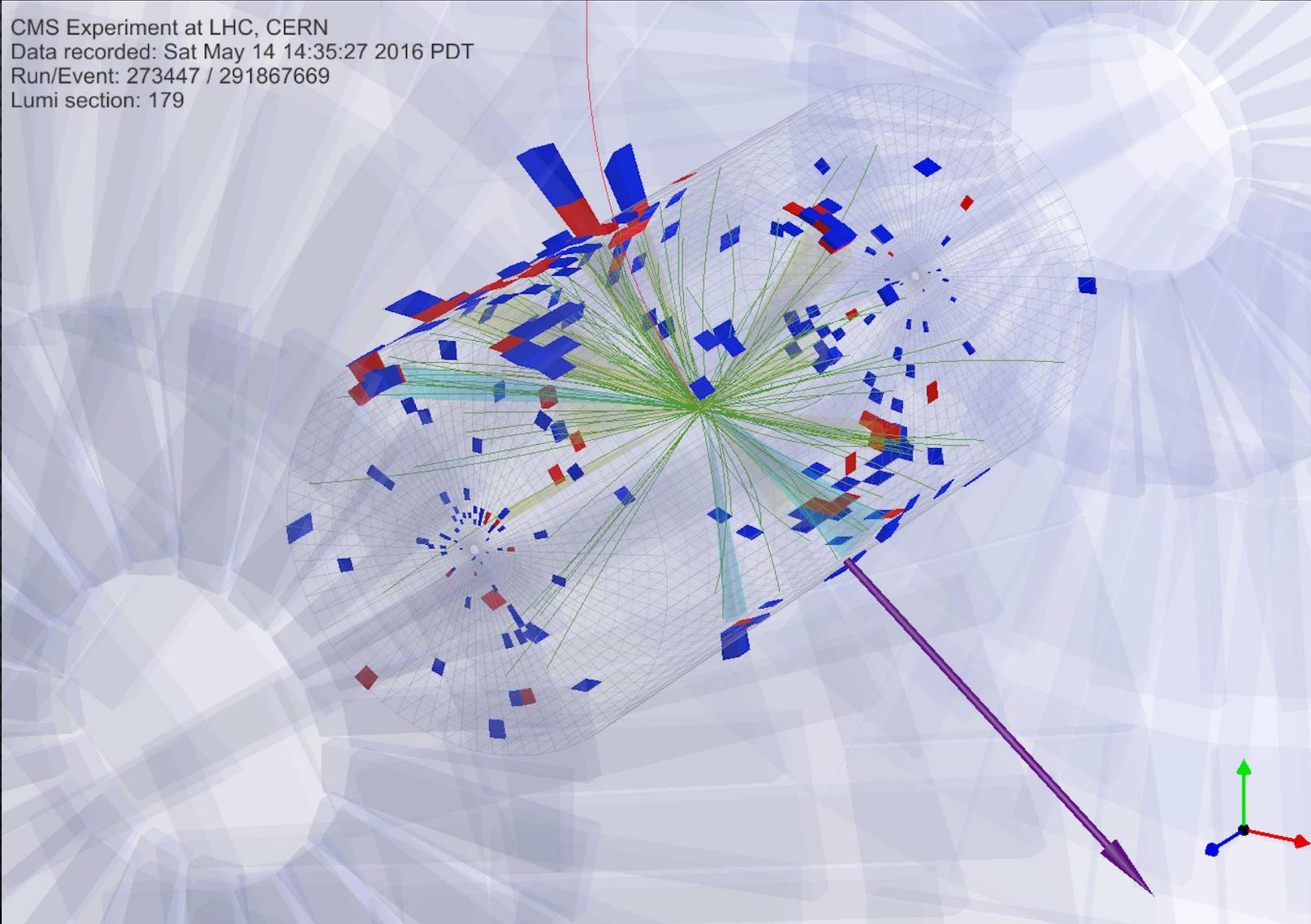
Can also use photon+jets data sample as the production of photons and Z bosons are very similar (at high energy where the mass of the Z is not relevant). Many more photon events than $Z \rightarrow \nu\nu$ so good statistical accuracy but need to carefully control for production differences.

This analysis uses both methods to estimate the $Z \rightarrow \nu\nu$ background

Possible SUSY event

This event contains 12 jets, 3 b-jets, and large MHT.

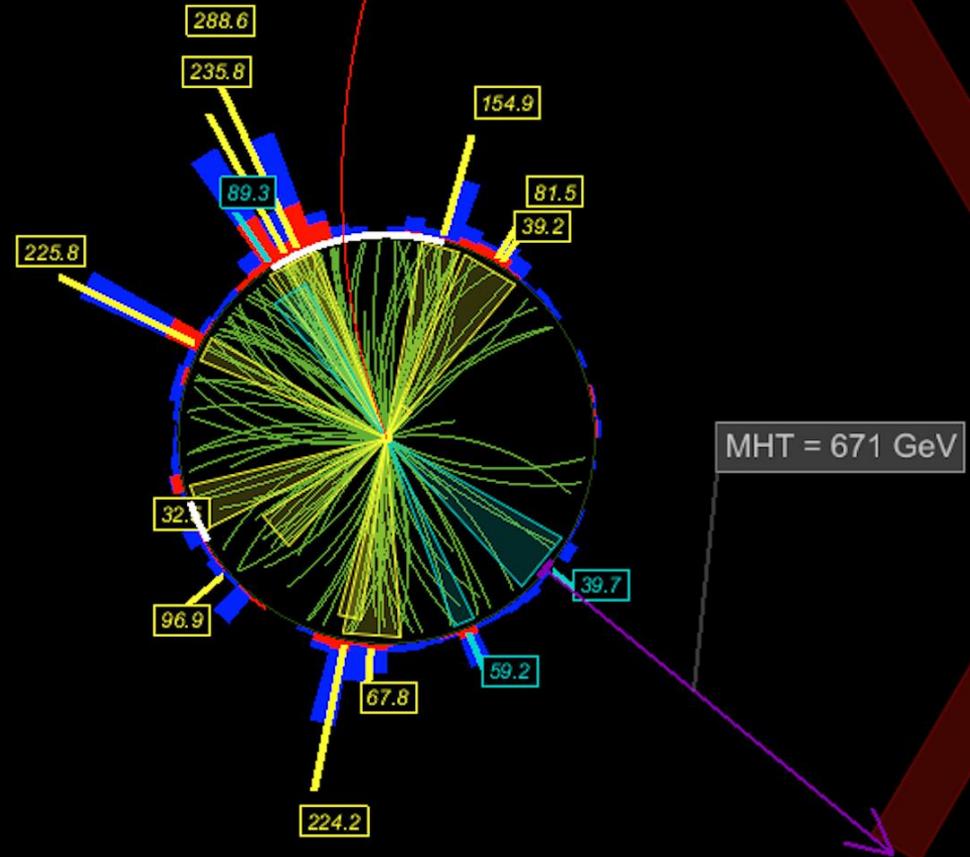
CMS Experiment at LHC, CERN
Data recorded: Sat May 14 14:35:27 2016 PDT
Run/Event: 273447 / 291867669
Lumi section: 179



Possible SUSY event

This event contains 12 jets, 3 b-jets, and large amount of missing transverse energy. Great candidate for a SUSY event (but most likely not).

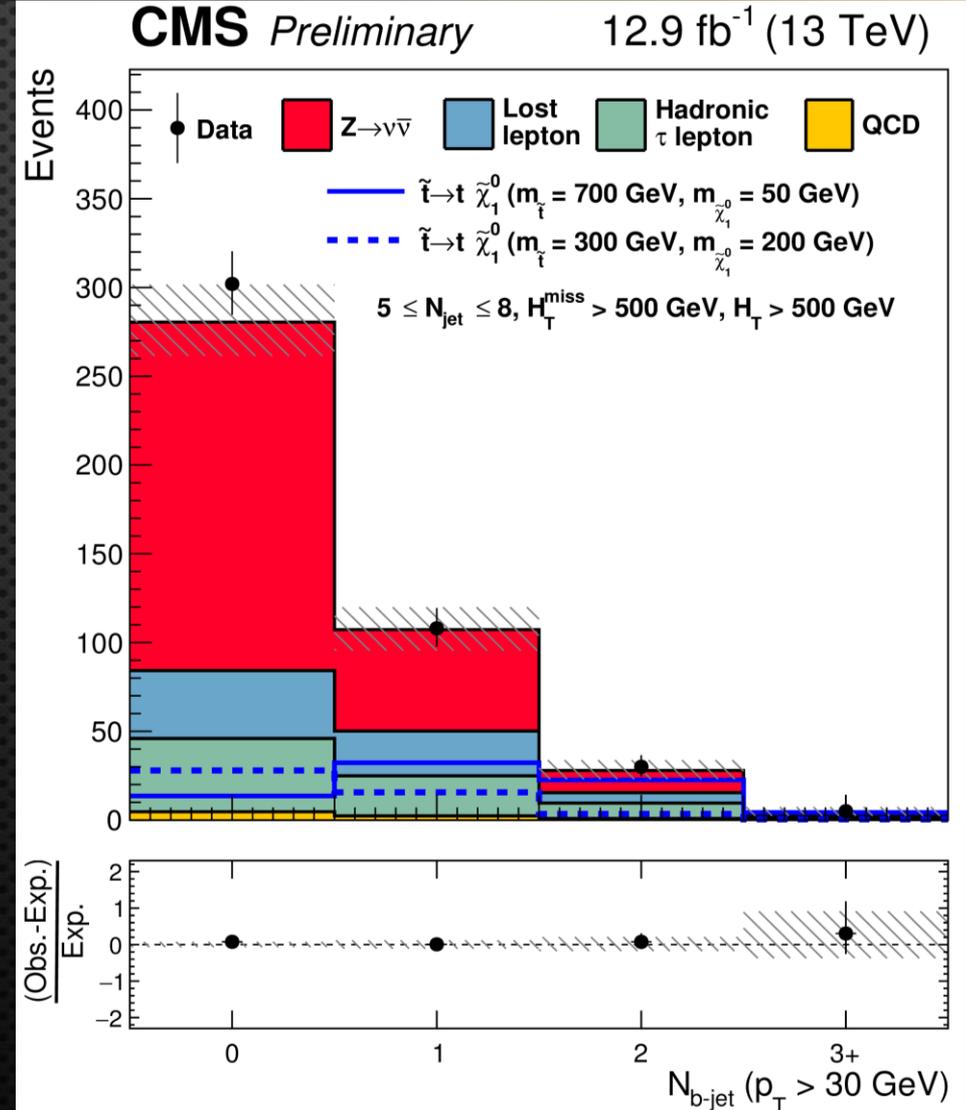
CMS Experiment at LHC, CERN
Data recorded: Sat May 14 14:35:27 2016 PDT
Run/Event: 273447 / 291867669
Lumi section: 179



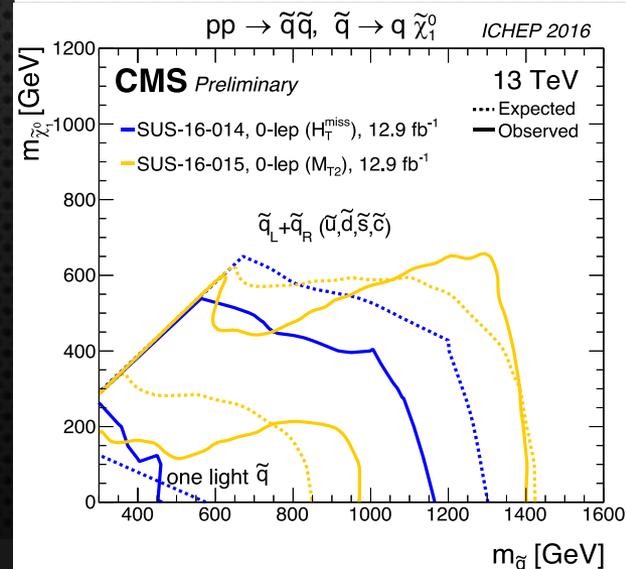
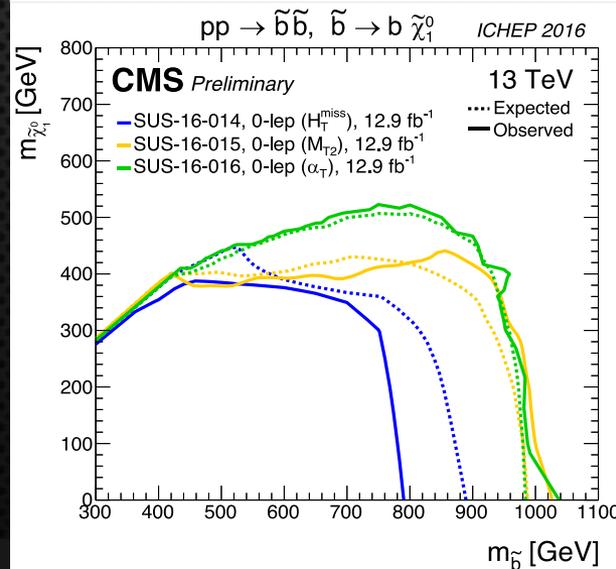
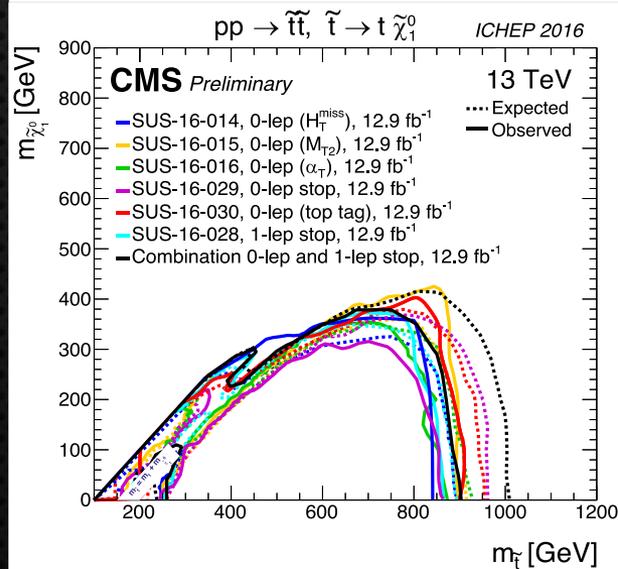
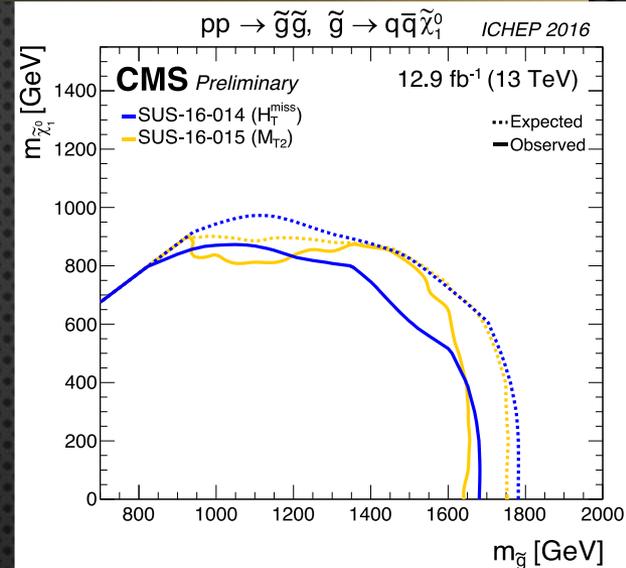
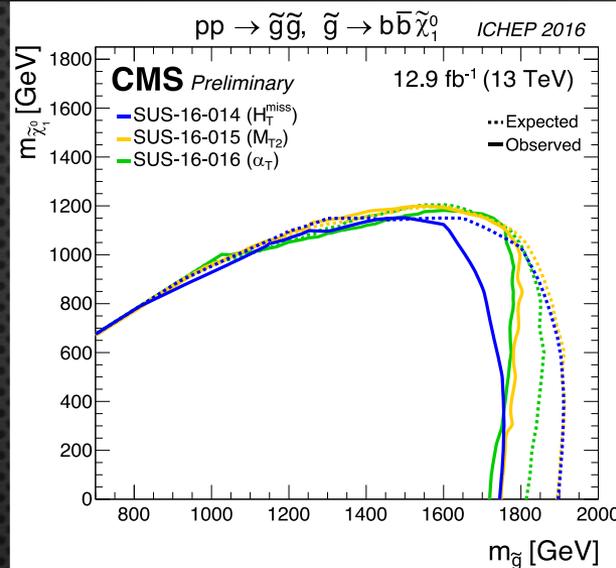
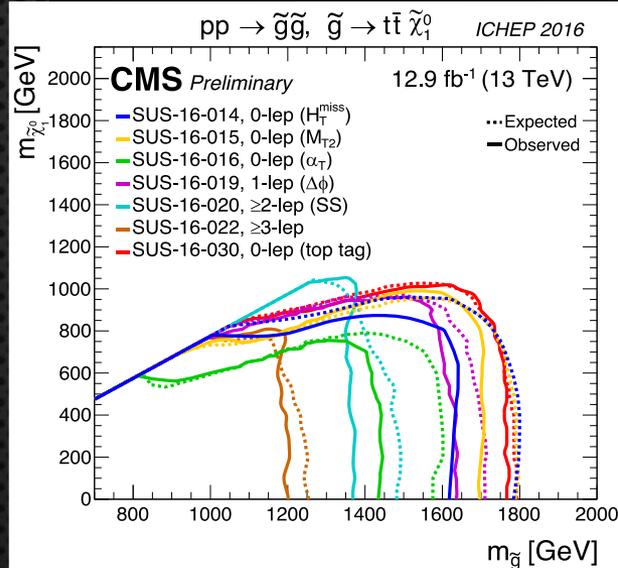
Search results

Can look in the more sensitive regions to see what a signal may look like.

Still no sign of SUSY.



Summary of hadronic SUSY searches



Other SUSY searches

- There are many other CMS searches for SUSY.
- These generally include some number of leptons ($e/\mu/\tau$) and/or target more specific models.
- Example of electroweak production (instead of strong production) with leptons in the final state.

