

"Spontaneous Breaking of Gauge Symmetry"

We can get a taste for the modern discussion of the photon mass (and related issues) if we diverge from pure electrodynamics to consider two related topics (which we will do by example):

- Goldstone's theorem
- the Higgs effect

Let's consider once again a classical field with n components

$$\mathcal{L} = \frac{1}{2} \sum_{j=1}^n (\partial_\mu \phi_j)(\partial^\mu \phi_j) - V(\phi)$$

where the "potential" is taken for simplicity to be

$$V(\phi) = \frac{\mu^2}{2} \phi^2 + \frac{\lambda}{4!} (\phi^2)^2$$

$$\text{and } \phi^2 = \sum_{j=1}^n \phi_j^2$$

These models arise in a variety of contexts

- In condensed matter physics ϕ might represent the average value of the spin of a patch of a system of atoms (in a magnet). The term $V(\phi)$ measures the energy of the patch due to self-interaction (does ϕ want to be large or small?). The $(\partial_\mu \phi)(\partial^\mu \phi)$ term measures the interaction between patches of spins, separated in space. ["Ginzberg-Landau model"]
- In macroscopic systems showing quantum behavior: as a description of the condensate in a Bose Einstein gas or in liquid Helium; here Ψ ($\equiv \phi$) is complex, the wavefn.
- As models for real fundamental particles - the Higgs

What does the potential term do $\stackrel{\text{to}}{\rightarrow}$ the field equations?

$$-\frac{\partial \mathcal{L}}{\partial \varphi} = \frac{\partial V}{\partial \varphi} \quad \text{provides a "force" which}$$

attempts to drive φ towards minima of V . A particularly interesting and simple case to study is the case when V has a minimum in φ , and we simply linearize the equations of motion (or expand \mathcal{L} quadratically) about that minimum. Then

$$V(\varphi) = V(\varphi_0) + \frac{1}{2} V''(\varphi_0) (\varphi - \varphi_0)^2$$

$$\mathcal{L} = \frac{1}{2} \partial_x \varphi \partial^x \varphi + \frac{1}{2} V''(\varphi_0) (\varphi - \varphi_0)^2 + \text{constant } V(\varphi_0)$$

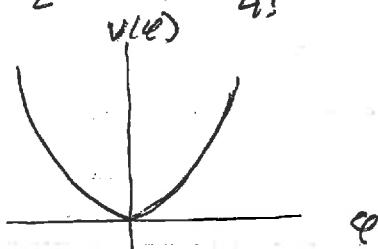
$$= \frac{1}{2} \partial_x (\varphi - \varphi_0) \partial^x (\varphi - \varphi_0) + \underbrace{\frac{1}{2} V''(\varphi_0)}_{\mu^2} (\varphi - \varphi_0)^2$$

\Rightarrow mass 2 = 2nd derivative of V at minimum.

- Hierarchy of examples -

example $\lambda = 1$ $V = \frac{1}{2} \mu_0^2 \varphi^2 + \frac{\lambda}{4!} \varphi^4$

a) $\mu_0^2 > 0$

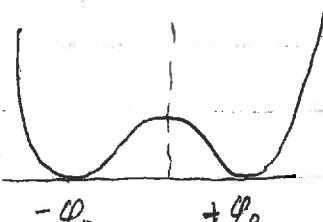


minimum is at $\varphi_0 = 0$

$$\mu^2 = \mu_0^2$$

b) $\mu_0^2 < 0$ & it's just a parameter

in a magnet we might parameterize it as $\mu^2 \propto T - T_c$
in a magnet $\mu_0^2 \propto T - T_c$

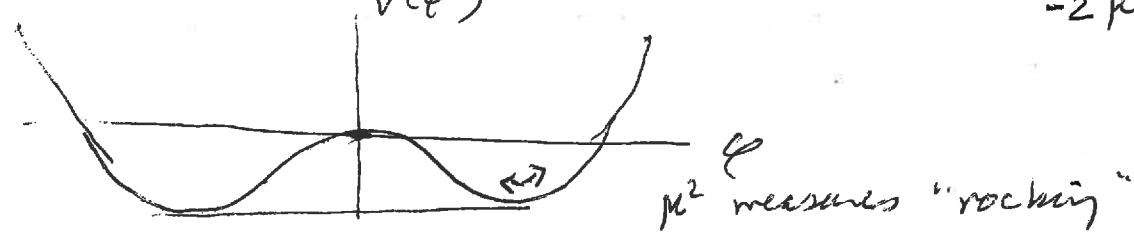


$$\left. \frac{\partial V}{\partial \varphi} \right|_{\varphi_0} = 0 = \mu_0^2 \varphi_0 + \frac{\lambda}{6} \varphi_0^3$$

$$\varphi_0^2 = -\frac{6\mu_0^2}{\lambda}$$

$$\frac{\partial^2 V}{\partial \varphi^2} = \frac{1}{2} \left[\mu_0^2 + \frac{\lambda}{2} \varphi_0^2 \right] = \frac{1}{2} (\mu_0^2 - 3\mu_0^2) \\ = -2\mu_0^2$$

(recall μ_0^2 was set < 0) \Rightarrow mass $^2 = \mu^2 = \frac{-2\mu_0^2}{\lambda}$



Notice that the original model was invariant under the discrete symmetry $\varphi(x,t) \rightarrow -\varphi(x,t)$.

If the system chooses to minimize its potential by sitting in one of the minima, we say the vacuum φ_0 symmetry is "broken" — and indeed if we write

$$\varphi(x,t) = \varphi_0 + \chi(x,t)$$

$$\mathcal{L} = \frac{1}{2} \partial_x \chi \partial^{\mu} \chi - V(\chi)$$

$$V(\chi) = \frac{1}{2} \mu_0^2 (\chi + \varphi_0)^2 + \frac{\lambda}{4!} (\chi + \varphi_0)^4 \\ = -\frac{1}{2} \mu_0^2 \chi^2 + \frac{\lambda}{6} \varphi_0 \chi^3 + \frac{\lambda}{4!} \chi^4 + \text{constants.}$$

(again $\mu_0^2 < 0$)

(see next page)

It's not obvious that $\chi \rightarrow -\varphi_0 - (\chi + \varphi_0)$ ($\varphi \rightarrow -\varphi$) is a symmetry. "An ant living in a magnetized magnet has a hard time realizing that the underlying system is rotationally invariant!"

Word "spontaneous symmetry breaking"

$$\begin{aligned}
 V(x+\phi_0) &= \frac{1}{2} \mu_0^2 [x + \phi_0]^2 + \frac{\lambda}{24} [x + \phi_0]^4 \\
 &= \frac{1}{2} \mu_0^2 [x^2 + 2x\phi_0 + \dots] \\
 &\quad + \frac{\lambda}{24} [x^4 + 4x\phi_0^3 + 6x^2\phi_0^2 + \dots] \\
 &= x^2 \left[\frac{1}{2} \mu_0^2 + \frac{\lambda}{4} \phi_0^2 \right] \quad x^2 \left(\frac{1}{2} \mu_0^2 - \frac{6}{4} \mu_0^2 \right. \\
 &\quad \left. + x \left[\mu_0^2 \phi_0 + \frac{4}{6} \phi_0^3 \right] \rightarrow 0 \right) = x^2 \mu_0^2 \\
 &\quad + \text{higher order}
 \end{aligned}$$

$$\phi_0 = -\frac{6\mu_0^2}{\lambda}$$

2nd example: $d=2$, or \mathbf{q} is 2-dimensional

$$V(\mathbf{q}) = \frac{1}{2} \mu_0^2 [q_1^2 + q_2^2] + \frac{\lambda}{4!} [q_1^2 + q_2^2]^2$$

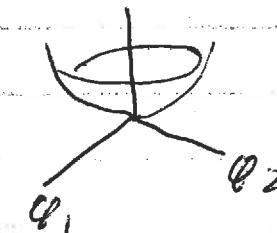
Note the symmetry

$$\begin{pmatrix} q_1' \\ q_2' \end{pmatrix} = \begin{bmatrix} \cos \varphi & \sin \varphi \\ -\sin \varphi & \cos \varphi \end{bmatrix} \begin{pmatrix} q_1 \\ q_2 \end{pmatrix}$$

or $\mathbf{q}' = R \mathbf{q}$. A global rotation of (q_1, q_2) leaves \mathcal{L} invariant. There is an associated conserved Noether current, as we found earlier.

~~Now if $\mu_0^2 > 0$ the potential surface looks a saddle.~~

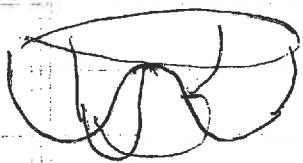
If $\mu_0^2 > 0$ the potential is concave up at a minimum at $q_1 = q_2 = 0$



$$\frac{1}{2} \left. \frac{\partial^2 V}{\partial q_i^2} \right|_{q_1=q_2=0} = \frac{1}{2} \left. \frac{\partial^2 V}{\partial q_2^2} \right|_{q_1=q_2=0} = 0$$

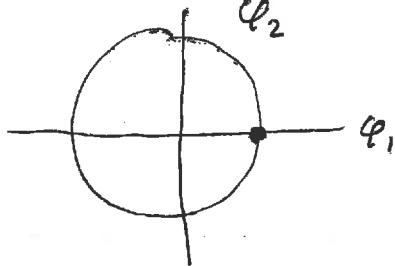
$$\text{and } \left. \frac{1}{2} \frac{\partial^2 V}{\partial q_1 \partial q_2} \right|_{q_1=q_2} = 0$$

But if $\mu_0^2 < 0$, the potential surface looks like sombrero or the bottom of a wine bottle



Defining $q_1^2 + q_2^2 = C^2$, the potential has a minimum at any point on a circle $C^2 = -\frac{6\mu_0^2}{\lambda^2}$

Let's arbitrarily suppose that the vacuum chooses to break the symmetry by setting $\varphi_1 = \sqrt{\frac{6\mu_0^2}{\lambda}}, \varphi_2 = 0$



$$\begin{aligned} &= \varphi_0 \\ &\varphi_0^2 = -\frac{6\mu_0^2}{\lambda} \\ &(\text{recall } \mu_0^2 < 0) \end{aligned}$$

What is the spectrum?

Write $\varphi_i = \varphi_0 + \chi_i$?/?
 $\varphi_2 = \chi_2$

$$\begin{aligned} V = & \frac{1}{2} \mu_0^2 \left[(\varphi_0 + \chi_1)^2 + \chi_2^2 \right] + \frac{\lambda}{24} \left[(\varphi_0 + \chi_1)^2 + \chi_2^2 \right]^2 \\ = & \frac{1}{2} \mu_0^2 \left[\varphi_0^2 + 2\varphi_0 \chi_1 + \chi_1^2 + \chi_2^2 \right] \quad \text{wrt page} \\ & + \frac{\lambda}{24} \left[\varphi_0^2 + 2\varphi_0 \chi_1 + \chi_1^2 + \chi_2^2 \right]^2 \\ = & \chi_1 \left[+\varphi_0 \mu_0^2 + \frac{\lambda}{24} \varphi_0^3 \right] \left\{ \begin{array}{l} \leftarrow \text{zero: } \mu_0^2 + \frac{\lambda \varphi_0^2}{6} = 0 \\ \leftarrow \text{nonzero } \frac{1}{2} \mu_0^2 + \frac{\lambda}{4} \varphi_0^2 = -\frac{3}{2} \mu_0^2 \end{array} \right. \\ & + \chi_1^2 \left[\frac{1}{2} \mu_0^2 + \frac{\lambda}{24} (4\varphi_0^2 + 2\varphi_0^3) \right] \\ & + \chi_2^2 \left[\frac{1}{2} \mu_0^2 + 2 \frac{\lambda}{24} \varphi_0^2 \right] \left\{ \begin{array}{l} \leftarrow \text{zero!} \\ = \frac{1}{2} \mu_0^2 - \frac{3}{2} \mu_0^2 \\ = -\frac{1}{2} \mu_0^2 \end{array} \right. \\ & + \chi_1 \chi_2 [0] \\ & + \text{higher order terms} \end{aligned}$$

Quadratic part = $\underbrace{(-\mu_0^2)}_{0 \gg \frac{m^2}{2}} \chi_1^2 + 0 \cdot \chi_2^2$

The χ_2 mode is massless. This is a general feature.

While the χ_1 mode is massive

$$\frac{1}{2} \mu_0^2 \left[(\varphi_0 + \chi_1)^2 + \chi_2^2 \right] + \frac{\lambda}{24} \left[(\varphi_0 + \chi_1)^2 + \chi_2^2 \right]^2$$

$$= \frac{1}{2} \mu_0^2 \left[\varphi_0^2 + 2\varphi_0 \chi_1 + \chi_1^2 + \chi_2^2 \right]$$

$$+ \frac{\lambda}{24} \left[\varphi_0^2 + 2\varphi_0 \chi_1 + \chi_1^2 + \chi_2^2 \right]^2$$

$$= \chi_1 \left[\varphi_0 \mu_0^2 + \frac{\lambda}{24} \cdot 4 \cancel{\varphi_0^2} 2 - 2\varphi_0 \cdot \varphi_0^2 \right]$$

$$+ \chi_1^2 \left[\frac{1}{2} \mu_0^2 + \frac{\lambda}{24} (2\varphi_0^2 + 4\varphi_0^2) \right]$$

$$+ \chi_2^2 \left[\frac{1}{2} \mu_0^2 + 2\frac{\lambda}{24} (\cancel{\varphi_0^2}) \right]$$

$$+ \chi_1 \chi_2 \cdot 0$$

+ higher order

$$\cancel{\varphi_0^2} \text{ Now } \varphi_0^2 + \cancel{\varphi_0^2} = -\frac{6\mu_0^2}{\lambda}$$

$$\chi_1 \varphi_0 (\mu_0^2 + 6\lambda \varphi_0^2) = \chi_1 \cdot 0$$

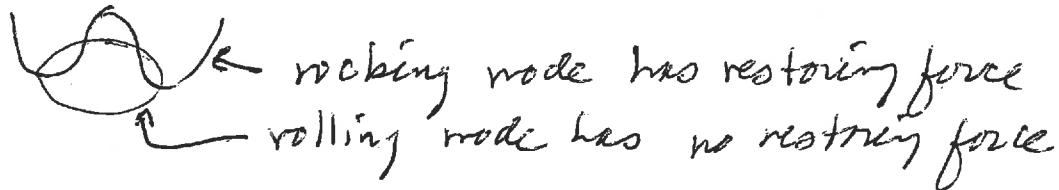
$$+ \chi_1^2 \left(\frac{1}{2} \mu_0^2 + \frac{\lambda}{4} \left(-\frac{6\mu_0^2}{\lambda} \right) \right) = -\frac{1}{2} \mu_0^2 \chi_1^2 \text{ (no sym)}$$

$$+ \chi_2^2 \left[\frac{1}{2} \mu_0^2 + \frac{\lambda}{12} \varphi_0^2 \right] = \chi_2^2 \cdot 0$$

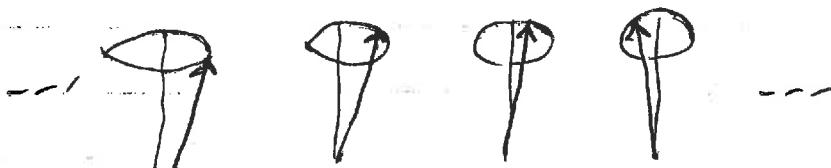
"When a global continuous symmetry is broken,
there is an accompanying massless mode" spontaneously

≡ Goldstone's theorem

$$\text{Massless} \equiv E(k) \propto k$$



Massless mode is called a "Goldstone Boson"
example: spin wave in a magnet-system
can support arbitrarily long-wavelength, low
energy excitations, with spins precessing around
the ordered direction, ~~like waves~~



This is $\varphi = (\varphi_1, \varphi_2, \varphi_3)$ with $\langle \varphi_3 \rangle = v$



Interesting ~~aspects~~ group theory ~~strong~~ stories ---

Higgs Effect.

This is still not a massive photon. Recall that electrodynamics appeared when we tried to enforce local phase rotations as a symmetry.

$$\phi(x) \rightarrow e^{iA(x)} \phi(x)$$

or $\begin{pmatrix} \phi_1(x) \\ \phi_2(x) \end{pmatrix} \rightarrow \begin{pmatrix} \cos\theta(x) & \sin\theta(x) \\ -\sin\theta(x) & \cos\theta(x) \end{pmatrix} \begin{pmatrix} \phi_1(x) \\ \phi_2(x) \end{pmatrix}$

Let's treat ϕ as complex, write

$$\mathcal{L} = [(\partial_\mu + ieA_\mu)\phi][(\partial^\mu + ieA^\mu)\phi^*] + \mu_0^2 \phi^* \phi - \lambda (\phi^* \phi)^2 - \frac{1}{4} F_{\mu\nu} F^{\mu\nu}$$

and ~~forget~~ look at the broken symmetry case.

Note: $+\mu_0^2$ corresponds to sombrero (to avoid unphysical factors and μ, λ) convention of Ryder, Quantum Field Theory p.301

$$\text{The minimum of } V \text{ is at } |\phi| = \sqrt{\frac{\mu_0^2}{2\lambda}} \equiv \frac{a}{\sqrt{2}}$$

$$\text{Now write } \phi(x) = \frac{a + X\phi_1(x) + iX\phi_2(x)}{\sqrt{2}}$$

i.e. move slightly off the ~~minimizing~~ circle and re-express in terms of the physical fields ϕ_1, ϕ_2 .

~~Re-insert; (aside) and look only at the quadratic terms~~

$$\begin{aligned} \mathcal{L} = & -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + e^2 a^2 A_\mu A^\mu + \frac{1}{2} (\partial_\mu \phi_1)^2 + \frac{1}{2} (\partial_\mu \phi_2)^2 \\ & - 2\lambda a \phi_1^2 + \sqrt{2} ea A_\mu \partial_\mu \phi_2 + \text{cubic} + \text{quartic} \\ & \phi^* \phi_1 \phi_2 \end{aligned}$$

Let's check what we have

a) potential term

$$V(\phi) = -\mu_0^2 \left| \frac{a + X_1 + iX_2}{\sqrt{2}} \right|^2 + \frac{\lambda}{4} \left(|a + X_1 + iX_2|^2 \right)^2$$

$$= -\frac{\mu_0^2}{2} \left[a^2 + 2aX_1 + X_1^2 + X_2^2 \right]$$

$$+ \frac{\lambda}{4} \left(a^2 + 2aX_1 + X_1^2 + X_2^2 \right)^2$$

expand, keep linear + quadratic, discard the rest

$$V(\phi) = X_1 \left[-a\mu_0^2 + 4a^3 \frac{\lambda}{4} \right] \xrightarrow{a^2 = \mu_0^2 \lambda} 0$$

$$+ X_2^2 \left[-\frac{\mu_0^2}{2} + \cancel{\frac{(4a^2 + 2a^2)\lambda}{4}} \right] \xrightarrow{-\frac{\mu_0^2}{2} + \frac{3}{2}\lambda a^2} = \frac{\lambda a^2}{2}$$

$$+ X_1 X_2 \cdot 0$$

$$+ X_2^2 \cdot \left[-\frac{\mu_0^2}{2} + 2a^2 \frac{\lambda}{4} \right] \xrightarrow{0} 0$$

$$V(\phi) = \frac{\lambda a^2}{2} X_2^2 + 0 \cdot X_2^2$$



$$D_\mu = \partial_\mu - ieA_\mu \rightarrow \varphi = \frac{a + X_1 + iX_2}{\sqrt{2}}$$

$$D_\mu \varphi = \frac{1}{\sqrt{2}} \left[\partial_\mu X_1 - ieA_\mu X_1 - ieA_\mu a + ie\partial_\mu X_2 - eA_\mu X_2 \right]$$

$$(D_\mu \varphi)^* (D^\mu \varphi) = \frac{1}{2} \left[\partial_\mu X_1 \partial^\mu X_1 + \partial_\mu X_2 \partial^\mu X_2 \right]$$

$$+ \frac{1}{2} e^2 a^2 A_\mu A^\mu$$

$$+ \frac{2}{\sqrt{2}} ea A_\mu \partial^\mu X_2$$

+ non quadratic

If $a m - A_\mu A^\mu -$ photon seems to have acquired a mass! Also X_1 is massive.

But - what is the funny $A_\mu \partial^\mu X_2$ term?

To make the answer less confusing, make a change of variables - a local gauge transformation

$$\phi' = \exp[iA(x)] \cdot \phi \quad A'_\mu = A_\mu + \frac{1}{e} \partial_\mu A$$

$$\phi' \sim [1+iA] \phi = (1+iA) \left[\frac{\alpha + X_1 + iX_2}{\sqrt{2}} \right]$$

$$= \frac{\alpha + (X_1 - A X_2) + i(X_2 + A X_1)}{\sqrt{2}}$$

$$= \frac{\alpha + X'_1 + iX'_2}{\sqrt{2}}$$

Choose A so that $X'_2 = 0$

In that gauge

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu}^i F^{i\mu\nu} + \frac{1}{2} e^2 a^2 A'_\mu A'^\mu$$

$$+ \frac{1}{2} (D_\mu X_1)^2 - \frac{\lambda a^2}{2} X_1^2$$

+ non-quadratic

Relabel A' as A , X as X . We have a Lagrangian

for A_μ - massive gauge boson, $M_A^2 \sim e^2 a^2 = \frac{e^2 g^2}{\lambda}$

X_1 massive scalar boson - the Higgs field

or particle - $M_H^2 = \lambda a^2$

This procedure is called the Higgs effect

Contrast

Goldstone mode: spontaneous break of global $U(1)$)

2 massive scalar fields \Rightarrow 1 massive scalar
 when symmetry unbroken
 1 massless scalar
 when broken

Higgs mode (SB of local $U(1)$)

2 massive scalars
 $+ \qquad \qquad \Rightarrow \qquad \qquad$ 1 massive scalar
 1 massless photon $\qquad \qquad \qquad$ 1 massive photon
 $\qquad \qquad \qquad$ (3 pol.-states)

$2 + 1 \cdot 2$ helicity = 4 modes states of photon = $1 + 1 \cdot 3 m_J$'s = 4 modes

"the photon has eaten the scalar mode and
 acquired a mass"

N.B. The original gauge invariance is still a symmetry
 of the Lagrangian, though not of the ~~vacuum~~
~~exterior~~ vacuum - there is still a conserved current.

Examples: ① Meissner effect in superconductor

see S. Weinberg Prog. Theor. Phys. Suppl. 86 43 (1986)
 or Ryder p.305

② W and Z mass ($U(1) \rightarrow SU(2) \times U(1)$)

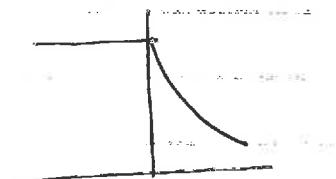
More complicated group theory: 4 higgs,

3 eaten $\xrightarrow{\text{GBS}}$ m_{W^\pm}, m_Z

1 left as physical massive higgs - ~~were~~ ~~now~~
 the particle at 125 GeV

Meissner effect

Superconductivity is always accompanied by Meissner effect: $E + H$ fields vanish inside sc. exponentially



$(O = \infty)$. Exponential suppression is basically mass generation for photon.

Below T_c , electron-phonon interactions (int with lattice vibrations) lead to effective attractive int. of electrons - system can lower its energy from usual free electron gas by forming bound state (Cooper pairs)

$\Psi_s(r)$ = wavefn of SC state

$$|\Psi_s(r)|^2 = n_{\text{pairs}}(r) = \frac{n_s(r)}{2} \text{ for electrons in SC state}$$

at $T > T_c$, $n_s = 0$. Write a free energy function of SC in terms of Ψ_s , free energy density is function of Ψ_s , $\nabla \Psi_s$. Ψ_s is a charged field

(with charge $e^* = 2e$, mass $m^* = 2m$) so

F depends on $|\Psi|^2$ and canonical momentum

$$\frac{\hbar}{i} \nabla \rightarrow \frac{\hbar}{i} \nabla - e^* \frac{A}{c}$$

$$F(\Psi_s) = F_{\text{normal}}(T, 0) + \int d^3r \frac{B(r)^2}{8\pi}$$

$$+ \int d^3r [a |\Psi_s|^2 + \frac{b}{2} |\Psi_s|^4 + \dots]$$

$$+ \int d^3r \left[\frac{1}{2m^*} \left(\frac{\hbar}{i} \vec{\nabla} - \frac{e^*}{c} \vec{A}(r) \right) \vec{\Psi}_s(r) \right]^2 + \dots$$

and $B = \nabla \times A$. We want to minimize F by varying Ψ . This is a "typical" variational problem - it should be no surprise that the min of F happens if eqns of motion are satisfied, namely

$$\vec{\nabla} \times \vec{B} = \frac{4\pi}{c} \vec{J}_s \quad \text{where} \quad (1)$$

$$\vec{J}_s = \frac{e^* h}{2m^* c} \left(\psi_s^* \vec{\nabla} \psi - (\vec{\nabla} \psi_s) \vec{\nabla} \psi \right) \quad (2)$$

$$- \frac{e^2}{m^* c} |\psi_s|^2 \vec{A}(r) \quad \text{diff } F_{\text{elect}}$$

$$\left\{ \frac{1}{2m^*} \left(\frac{h}{c} \vec{\nabla} - \frac{e^* A}{c} \right)^2 + b |\psi_s|^2 \right\} \psi_s = -a \psi_s \quad (3)$$

Notice ~~that~~ the current: It comes from

$$\frac{\partial \Omega}{\partial A_\mu} = -\vec{J}_\mu$$

where Ω is the free energy density = (1) + (3) are

$$\Omega = \frac{\partial \Omega}{\partial \mu} \frac{\partial \Omega}{\partial \mu A_\mu} - \frac{\partial \Omega}{\partial A_\mu} = 0$$

in (3), $-a$ is "like" an energy eigenvalue and the $b |\psi_s|^2$ is "like" a repulsive self interaction which tries to spread Ψ_s over the whole volume.

decs

Now for solutions inside the SC, $\vec{B} = 0$, n_s is a constant. Then the free energy shift from the normal state is

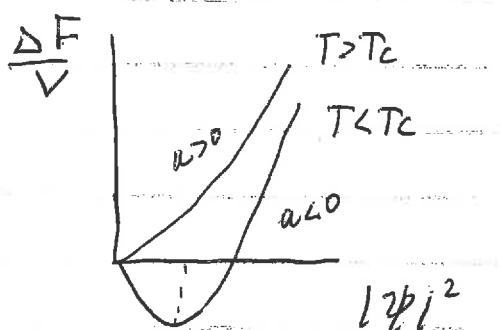
$$\Delta F = F_s - F_N = V \cdot \left(a |\psi_s|^2 + \frac{b}{2} |\psi_s|^4 \right)$$

with $(a + b|\psi_s|^2) \Psi_s = 0$

$b > 0$ so if $a > 0$, then $\Psi_s = 0$, the system is in its normal state

but if $T < T_c$, to get a nontrivial ground state,

$a < 0$ —



$$|\psi_s|^2 = \frac{n_s}{2} = -\frac{a}{b} > 0$$

$$\frac{\Delta F}{V} = -\frac{a^2}{2b} \quad a < 0$$

~~Discuss~~ Now imagine adding a small magnetic field \vec{B} . The current is (ψ_s is constant in space)

$$\vec{J}_s = -\frac{e^2}{m^* c} |\psi_s|^2 \vec{A} = -\frac{e^2 n_s}{m^* c} \vec{A}$$

~~Discuss~~ so $\vec{\nabla} \times \vec{J} = -\frac{e^2 n_s}{m^* c} \vec{B}$

Now $\vec{\nabla} \times \vec{B} = \frac{4\pi}{c} \vec{J}$, take curl again

$$\vec{\nabla} \times (\vec{\nabla} \times \vec{B}) = \vec{\nabla} (\vec{\nabla} \cdot \vec{B}) - \vec{\nabla}^2 \vec{B} = \frac{4\pi}{c} \vec{\nabla} \times \vec{J} = -\frac{4\pi n_s c^2}{m^* c^2} \vec{B}$$

$$\vec{\nabla}^2 \vec{B} = \frac{1}{c^2} \vec{B}, \quad \lambda = \sqrt{\frac{m^* c^2}{4\pi n_s c^2}}$$

or - a magnetic field is screened.

λ is called the "London penetration depth"
but you can see it's basically a photon mess.