

## Noether's theorem

Classical mechanics for a set of classical variables  $q_i$  &  $\dot{q}_i$  defined for all  $t$  begins with ~~an action~~ a Lagrangian  $L$  and an action  $S$

$$S = \int dt L(q_i, \dot{q}_i)$$

and  $\delta S = 0 \Rightarrow \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_i} - \frac{\partial L}{\partial q_i} = 0$

Classical mechanics for ~~fields~~ classical fields begins with the fields themselves, defined for all points in space  $x$  and time  $t$  - generically  $\phi(x, t)$  and their derivatives  $\frac{\partial \phi}{\partial x_j}, \frac{\partial \phi}{\partial t}$ ; generically,  $\partial_\mu \phi$ .

The analog of  $L$  is the Lagrange density  $\mathcal{L}$

$$L = \int \mathcal{L} d^3x$$

and the action is  $\int dt \cdot \int d^3x = \int d^4x$

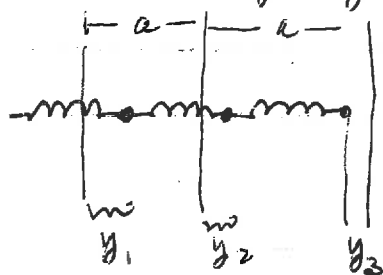
$$S = \int d^4x \mathcal{L}(\phi, \partial_\mu \phi)$$

The bottom line, if  $\delta S = 0$ :

$$\partial_\mu \left[ \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \right] - \frac{\partial \mathcal{L}}{\partial \phi} = 0$$

is our starting point. Let's do an example to refresh memory:

1-d Lagrangian for a set of mass points, springs



$y_j = y(x_j, t)$  = displacement of  $j$ th mass from equilibrium

$$L = \sum_{j=1}^N \frac{m}{2} \dot{y}_j^2 - \frac{1}{2} k (y_j - y_{j+1})^2 - \tilde{V}(y_j)$$

eqn of motion;  $m \ddot{y}_j + k [y_{j-1} - 2y_j + y_{j+1}] - \frac{\partial \tilde{V}}{\partial y_j} = 0$

$\left( \frac{d}{dt} \frac{\partial L}{\partial \dot{y}_j} - \frac{\partial L}{\partial y_j} = 0 \right)$

Now suppose  $y$  varies smoothly with  $x$ :

$$y_{j+1} = y_j + a \left. \frac{dy}{dx} \right|_{x=x_j} + \frac{a^2}{2} \left. \frac{d^2 y}{dx^2} \right|_{x=x_j} + \dots$$

$$m \ddot{y} + k a^2 \frac{d^2 y}{dx^2} + \frac{\partial \tilde{V}}{\partial y_j} = 0$$

$$= \frac{a}{k a^2} \left[ \frac{m}{k a^2} \frac{d^2 y}{dt^2} - \frac{d^2 y}{dx^2} - \frac{\partial \tilde{V}}{\partial \left( \frac{y}{k a^2} \right)} \right] \frac{1}{k a^2}$$

$$\phi(x, t) = \frac{y}{k a^2} \Rightarrow \frac{1}{c^2} \equiv \frac{m}{k a^2} \quad \text{with } V = \frac{\tilde{V}}{k a^2}$$

$$\frac{1}{c^2} \frac{\partial^2 \phi}{\partial t^2} - \frac{\partial^2 \phi}{\partial x^2} - \frac{\partial V}{\partial \phi} = 0$$

$$\partial_\mu \partial^\mu \phi - \frac{\partial V}{\partial \phi} = 0$$

$$= \partial_\mu \left[ \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi)} \right] - \frac{\partial \mathcal{L}}{\partial \phi}, \quad \mathcal{L} = \frac{1}{2} (\partial_\mu \phi) (\partial^\mu \phi) - V(\phi)$$

Now back to  $\mathcal{L}$ . Suppose we imagine only a change in the field variables

$$\varphi(x,t) \rightarrow \varphi(x,t) + \delta\varphi(x,t) \quad (1)$$

and at the same time

$$\begin{aligned} \partial_\mu \varphi(x,t) &\rightarrow \partial_\mu \varphi(x,t) + \partial_\mu \delta\varphi(x,t) \\ &= \partial_\mu \varphi(x,t) + \delta[\partial_\mu \varphi(x,t)] \end{aligned}$$

~~These are called "global" symmetry transformations when  $\delta\varphi$  is independent of  $x,t$ .~~

The change in  $\mathcal{L}$  is

$$\delta\mathcal{L} = \int_V \left[ \frac{\partial\mathcal{L}}{\partial\varphi_j} \delta\varphi_j + \frac{\partial\mathcal{L}}{\partial(\partial_\mu\varphi_j)} \delta(\partial_\mu\varphi_j) \right]$$

Recall the equations of motion  $\frac{\partial\mathcal{L}}{\partial\varphi_j} = \partial_\mu \left[ \frac{\partial\mathcal{L}}{\partial(\partial_\mu\varphi_j)} \right]$

$$\text{so } \delta\mathcal{L} = \int_V \left[ \partial_\mu \left( \frac{\partial\mathcal{L}}{\partial(\partial_\mu\varphi_j)} \right) \right] \delta\varphi_j + \frac{\partial\mathcal{L}}{\partial(\partial_\mu\varphi_j)} \delta(\partial_\mu\varphi_j)$$

$$= \partial_\mu \left[ \frac{\partial\mathcal{L}}{\partial(\partial_\mu\varphi_j)} \int_V \frac{\partial\mathcal{L}}{\partial(\partial_\mu\varphi_j)} \delta\varphi_j \right]$$

$$\equiv \partial_\mu J^\mu \quad \text{defining } [ ] \text{ as a current.}$$

Now if it happens that  $\delta\mathcal{L} = 0$ , the Lagrangian is invariant under the change of  $\varphi$ 's

then 
$$\partial_\mu J^\mu = 0$$

That is,  $J^\mu$  is a conserved current.

Symmetry implies conservation law

The "charge" associated with the current is also conserved

$$Q \equiv \int d^3x J_0$$

Changes in  $Q$  could be "internal" - redefines  $Q$  at every space-time point.

ex:  $\varphi(x,t) \rightarrow e^{i\theta(x,t)} \varphi(x,t)$

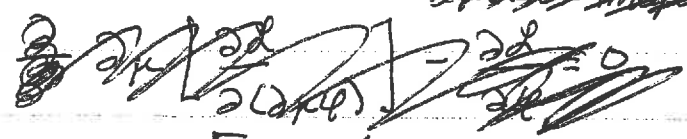
$\theta(x,t) = \text{constant}$ ; ~~glob~~ in  $x,t$   
example of a global symmetry transf.

$\theta(x,t)$  varies w/  $x,t$ : "local symmetry transformation"

or SC could involve changes in coordinates.

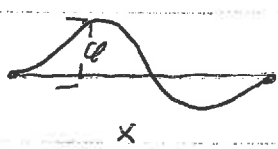
Later ones lead to "conserved currents" associated with energy & momentum.

Internal ones are more interesting for the moment



Example

Let us suppose for the sake of argument that we have a classical physical system described by a complex field  $\phi(x) = \phi_1(x) + i\phi_2(x)$  (or we could think of as a pair of real fields). A Boulder-style realization of  $\phi$  could be the QM wave function of a Bose-Einstein condensate, ~~or the order parameter~~ be it of atoms or of liquid helium. Suppose that its Lagrange density does not have any explicit dependence on space and time - instead, that  $\mathcal{L}$  only depends on  $x$  and  $t$  through  $\phi$  and  $\partial_x \phi$ . \* That would be the case for a stretched "string"



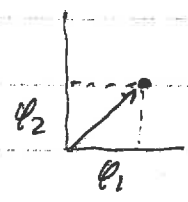
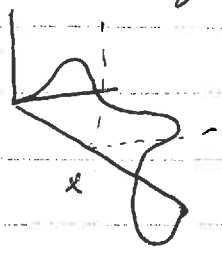
where  $\phi(x,t)$  is the local displacement of the string, and you recall  $[c=1]$

$$\mathcal{L} = \left(\frac{\partial \phi}{\partial t}\right)^2 - (\nabla \phi)^2$$

$$\equiv (\partial_x \phi)^2$$

A complex field would have a pair of  $\phi$ 's

so imagine  
m 1-d



stick to ~~(x,t)~~  
etc

~~TRP~~

\* turns out, to construct relativistically consistent theories, you have to do this

Now suppose that we imagine making a <sup>global</sup> ~~smooth~~ change ~~in the~~ rotation in the field variables

$$\begin{pmatrix} \psi'_1 \\ \psi'_2 \end{pmatrix} = \begin{bmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{bmatrix} \begin{pmatrix} \psi_1 \\ \psi_2 \end{pmatrix}$$

$$\text{or } \psi' = e^{i\theta} \psi \quad \text{if we call } \psi' \equiv (\psi_1 + i\psi_2)$$

and suppose that under this change the  $\mathcal{L}$  is invariant. For the case of the string, this is just a rotation in the field space  $(\psi_1, \psi_2)$ .

Let's suppose that  $\theta$  is infinitesimal,  $\theta = g\epsilon$

$$\psi' = \psi + \delta\psi$$

$$\delta\psi = i g \epsilon \psi$$

$\epsilon$  is the same everywhere in space. How does  $\mathcal{L}$  change?

~~$$\delta\mathcal{L} = \frac{\partial\mathcal{L}}{\partial\psi} \delta\psi + \frac{\partial\mathcal{L}}{\partial(\partial_x\psi)} \delta(\partial_x\psi)$$~~

~~but  $\delta\psi = i g \epsilon \psi$~~

~~$$\delta(\partial_x\psi) = i g \epsilon \partial_x\psi$$~~

~~$$\delta\mathcal{L} = i g \epsilon \left[ \frac{\partial\mathcal{L}}{\partial\psi} \psi + \frac{\partial\mathcal{L}}{\partial(\partial_x\psi)} \partial_x\psi \right]$$~~

But use the equation of motion

~~$$\frac{\partial\mathcal{L}}{\partial\psi} = \partial_x \left[ \frac{\partial\mathcal{L}}{\partial(\partial_x\psi)} \right]$$~~

and

$$\delta\mathcal{L} = i g \epsilon \left[ \partial_x \left[ \frac{\partial\mathcal{L}}{\partial(\partial_x\psi)} \right] \psi + \frac{\partial\mathcal{L}}{\partial(\partial_x\psi)} \partial_x\psi \right]$$

$$\mathcal{L} = \frac{1}{2} \left[ (\partial_\mu \phi_1)^2 + (\partial_\mu \phi_2)^2 \right] - V(\phi_1, \phi_2)$$

$$\begin{bmatrix} \phi_1' \\ \phi_2' \end{bmatrix} = \begin{bmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{bmatrix} \begin{bmatrix} \phi_1 \\ \phi_2 \end{bmatrix} \underset{\text{small } \theta}{\sim} \begin{bmatrix} 1 & \theta \\ -\theta & 1 \end{bmatrix} \begin{bmatrix} \phi_1 \\ \phi_2 \end{bmatrix}$$

$$\begin{aligned} \text{i.e. } \delta\phi_1 &= \theta\phi_2 & ; & \delta\partial_\mu\phi_1 = \theta\partial_\mu\phi_2 \\ \delta\phi_2 &= -\theta\phi_1 & & \delta\partial_\mu\phi_2 = -\theta\partial_\mu\phi_1 \end{aligned}$$

$$\delta\mathcal{L} = \underbrace{\frac{\partial\mathcal{L}}{\partial(\partial_\mu\phi_1)}}_{\text{kinetic}} \delta(\partial_\mu\phi_1) + \underbrace{\frac{\partial\mathcal{L}}{\partial(\partial_\mu\phi_2)}}_{\text{kinetic}} \delta(\partial_\mu\phi_2) + \underbrace{\frac{\partial\mathcal{L}}{\partial\phi_1} + \frac{\partial\mathcal{L}}{\partial\phi_2}}_{\text{potential}}$$

Suppose  $V(\phi_1, \phi_2)$  is a function only of  $\phi_1^2 + \phi_2^2$  - obviously, "potential" part of  $\mathcal{L}$  is unchanged under this transf- and last terms give zero. Kinetic term

$$\begin{aligned} \delta\mathcal{L} &= (\partial^\mu\phi_1) [\theta\partial_\mu\phi_2] + (\partial^\mu\phi_2) [-\theta\partial_\mu\phi_1] \\ &= 0, \text{ trivial!} \end{aligned}$$

Yes, it's a symmetry. What is the associated conserved current?

$$\begin{aligned} J^\mu &= \sum_j \frac{\partial\mathcal{L}}{\partial(\partial_\mu\phi_j)} \delta\phi_j = (\partial^\mu\phi_1) \theta\phi_2 + (\partial^\mu\phi_2) (-\theta\phi_1) \\ &= (\phi_2 \partial^\mu\phi_1 - \phi_1 \partial^\mu\phi_2) \text{ up to an overall constant.} \end{aligned}$$

A lot like ~~Schrodinger~~ Schrodinger eqn ...

So far we considered a variation of  $\phi$

$$\delta\phi = i\epsilon\phi$$

with  $\epsilon$  a constant over space, and we saw that if  $\mathcal{L}(\phi') = \mathcal{L}(\phi)$ , with  $\phi' = \phi + \delta\phi$ , then the system has a conserved current. What if we now let

$\epsilon$  vary from point to point in space? That is, suppose that we asked for the following transformation to be a symmetry?

$$\delta\phi(x) = i\epsilon(x)\phi(x) \quad \text{infinitesimal}$$

$$\text{or } \phi(x') = \exp(i\epsilon(x))\phi(x) \quad \text{finite}$$

and want  $\delta\mathcal{L} = 0$

We know that the variation in  $\mathcal{L}$  is still

$$\delta\mathcal{L} = \frac{\partial\mathcal{L}}{\partial\phi}\delta\phi + \frac{\partial\mathcal{L}}{\partial(\partial_\mu\phi)}\delta(\partial_\mu\phi)$$

$$\text{but now } \delta(\partial_\mu\phi) = \partial_\mu\delta\phi$$

$$= \underbrace{i\epsilon\partial_\mu\phi}_{\text{old term}} + \underbrace{i\phi\partial_\mu\epsilon}_{\text{new term!}}$$

The equations of motion

$$\partial_\mu \frac{\partial\mathcal{L}}{\partial(\partial_\mu\phi)} = \frac{\partial\mathcal{L}}{\partial\phi}$$

again allow us to write

$$\delta\mathcal{L} = \epsilon(x)\partial_\mu J^\mu(x) + \frac{\partial\mathcal{L}}{\partial(\partial_\mu\phi)} i\phi\partial_\mu\epsilon$$

with  $J^\mu = i\phi \frac{\partial\mathcal{L}}{\partial(\partial_\mu\phi)}$  is the same current we had before - it is conserved

$$\delta\mathcal{L} = \epsilon(x)\partial_\mu J^\mu(x) + J^\mu_0(x)\partial_\mu\epsilon(x)$$

What about the second term?



The only way to make it zero is to introduce new ~~the~~ fields into  $\mathcal{L}$ , whose variation cancels the  $\partial_\mu \epsilon$  term. The current is a 4-vector, so let's add in a vector field  $A_\mu(x)$ , whose variation is

$$\begin{aligned} \delta A_\mu(x) &= \partial_\mu \epsilon(x) \\ \text{i.e.} \quad \vec{A}(x)' &= \vec{A}(x) + \vec{\nabla} \epsilon \\ A_\nu(x)' &= A_\nu(x) + \frac{\partial \epsilon}{\partial x^\nu} \end{aligned}$$

$$\begin{aligned} \text{Now } \delta \mathcal{L}' &= \delta \mathcal{L}_{\text{matter}} + \frac{\partial \mathcal{L}}{\partial A_\nu} \delta A_\nu + \frac{\partial \mathcal{L}}{\partial (\partial_\nu A_\lambda)} \partial_\nu \delta A_\lambda \\ \delta \mathcal{L}' &= \delta \mathcal{L} + \frac{\partial \mathcal{L}}{\partial A_\nu} \partial_\nu \epsilon + \frac{\partial \mathcal{L}}{\partial (\partial_\nu A_\lambda)} \partial_\nu \partial_\lambda \epsilon \\ &\quad + (\partial_\lambda \epsilon) \partial_\lambda \end{aligned}$$

This is zero as long as

$$\frac{\partial \mathcal{L}}{\partial A_\nu} = -\partial_\nu \mathcal{L} \rightarrow \text{unique specification of coupling of new field to the currents}$$

$$\mathcal{L} = \mathcal{L} - \mathcal{J}_\mu A^\mu \quad \text{invariant}$$

$$\text{and } \frac{\partial \mathcal{L}}{\partial (\partial_\nu A_\lambda)} = -\frac{\partial \mathcal{L}}{\partial (\partial_\lambda A_\nu)} \rightarrow \text{so that the } \frac{\partial \mathcal{L}}{\partial \partial_\nu \partial_\lambda \epsilon} \text{ and } \frac{\partial \mathcal{L}}{\partial \partial_\lambda \partial_\nu \epsilon} \text{ bits cancel}$$

~~so~~ says  $\mathcal{L}$  involves  $\partial_\mu A_\nu - \partial_\nu A_\mu$ .

~~so~~

The new condition does not give conservation laws, but instead conditions on the couplings of fields.

1) Suppose a global symmetry

$$\delta\psi \sim i\epsilon\psi$$

Leads to invariant. There is a conserved current

$$\partial_\mu J^\mu = 0$$

2) replace global symmetry by local symmetry

$$\delta\psi(x,t) \sim i\epsilon(x,t)\psi(x,t)$$

Force  $\delta\mathcal{L} = 0$  no cov.

a) need new gauge field, a 4-vector  $A_\mu(x)$

Coupling b)  $\mathcal{L}_I = -j^\mu A_\mu$  - the same ~~is~~ conserved current as in (1)

c)  $\mathcal{L}$  involves  $\partial_\mu A_\nu - \partial_\nu A_\mu \equiv F_{\mu\nu}$ .

Symmetry transformation is

$$\phi'(x) = e^{i q \int E(x) dx} \phi(x)$$

$$A'_\mu(x) = A_\mu + \partial_\mu E(x)$$

invariance of

A way to encode this transformation in  $\mathcal{L}$  is to replace the usual derivative

$\partial_\mu \phi$   
by the covariant derivative

$$D_\mu \phi = \partial_\mu \phi - i q A_\mu \phi$$

and to write  $\mathcal{L}$  as a function of  $(D_\mu \phi)^\dagger \phi$   
not  $\partial_\mu \phi \partial^\mu \phi$

$$\begin{aligned} D'_\mu \phi' &= \partial_\mu \phi' - i q A'_\mu \phi' = \partial_\mu [e^{i q E} \phi] \\ &\quad - i q [A_\mu + \partial_\mu E] e^{i q E} \phi \\ &= e^{i q E} (\partial_\mu \phi + i q \phi \partial_\mu E) \\ &\quad - i q A_\mu \phi e^{i q E} - i q \phi \partial_\mu E e^{i q E} \\ &= e^{i q E} [\partial_\mu \phi - i q A_\mu \phi] = e^{i q E} D_\mu \phi \end{aligned}$$

$$\begin{aligned} \mathcal{L} &= (D'_\mu \phi')^\dagger (D'_\mu \phi) - V(|\phi'|^2) \\ &= e^{-i q E} e^{i q E} (D_\mu \phi)^\dagger (D_\mu \phi) - V(|\phi|^2) \end{aligned}$$

is invariant

$$|\phi'|^2 = |e^{i q E} \phi|^2 \text{ of course}$$

And finally, what would be a good candidate  $\mathcal{L}$  for the new electromagnetic degrees of freedom?

To get a scalar, contract indices  
world index of tensors

$$\mathcal{L} = c_1 [F_{\mu\nu}^2] + c_2 [F_{\mu\nu}]^2 + \dots$$

$$F_{\mu\nu}^2 = F_{\mu\nu} F^{\mu\nu}$$

and indeed, the Lagrange density which produces ~~Maxwell's~~ Maxwell's equations is "the simplest thing you can write down" ☺

$$\mathcal{L} = -\frac{1}{16\pi} F_{\mu\nu} F^{\mu\nu} - \frac{1}{c} \mathbf{J} \cdot \mathbf{A} + \text{terms independent of } \mathbf{A}$$

obviously rel. covariant made of 4 vectors by construction  
The coefficients and signs are chosen arbitrarily so that the field equations have CGS conventions.  
- Why isn't there more? Let's defer that for a while  
lets first show how this  $\mathcal{L}$  produces Maxwell's equations as its equations of motion.

Write

$$\mathcal{L} = -\frac{1}{16\pi} g_{\lambda\kappa} g_{\nu\sigma} [\partial^\kappa A^\sigma - \partial^\sigma A^\kappa] [\partial^\lambda A^\nu - \partial^\nu A^\lambda] - \frac{1}{c} J_\alpha A^\alpha$$

$$\frac{\partial \mathcal{L}}{\partial(\partial^\beta A^\alpha)} = -\frac{1}{16\pi} g_{\lambda\kappa} g_{\nu\sigma} \left[ \delta_\beta^\kappa \delta_\alpha^\sigma F^{\lambda\nu} - \delta_\beta^\sigma \delta_\alpha^\kappa F^{\lambda\nu} + \delta_\beta^\lambda \delta_\alpha^\nu F^{\kappa\sigma} - \delta_\beta^\nu \delta_\alpha^\lambda F^{\kappa\sigma} \right]$$

$$= -\frac{4}{16\pi} F_{\beta\alpha} = -\frac{1}{4\pi} F_{\beta\alpha}$$

$$\text{we need } \partial^\beta \frac{\partial \mathcal{L}}{\partial(\partial^\beta A^\alpha)} - \frac{\partial \mathcal{L}}{\partial A^\alpha} = 0$$

The ~~transformation~~ transformation

$$A_\mu \rightarrow A_\mu + \partial_\mu \epsilon(x)$$

(x)

is called a local gauge transformation. Theories which are invariant under ~~gauge~~ local gauge transformations are called 'gauge theories' and include Electrodynamics

and  $\mathcal{L}$  actually under a matrix generalization of

$$\psi(x) \rightarrow e^{i g \epsilon(x)} \psi(x) \Rightarrow \psi \rightarrow R \psi$$

QCD (strong interactions)

Weinberg-Salam model (weak + electromagnetic

interactions)

i.e. all nature  $\mathcal{L}$  gauge invariant for GR = coordinate reparametrization

How to satisfy local gauge invariance?

$$\mathcal{L} = \mathcal{L}(\psi, D_\mu \psi, F_{\mu\nu})$$

~~step~~

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$$

$$D_\mu \psi = \partial_\mu \psi - i g A_\mu \psi \quad \text{covariant derivative}$$

for then

$$\frac{\partial \mathcal{L}}{\partial A_\mu} = \frac{\partial \mathcal{L}}{\partial (D_\mu \psi)} (-i g \psi)$$

$$= \frac{\partial \mathcal{L}}{\partial (\partial_\mu \psi)} (-i g \psi) = -J_\mu$$

$$\delta \mathcal{L} = \int d^4x \left( \frac{\partial \mathcal{L}}{\partial A_\mu} \delta A_\mu + \frac{\partial \mathcal{L}}{\partial (D_\mu \psi)} \delta (D_\mu \psi) \right) = 0$$

$$\text{And } \frac{\partial \mathcal{L}}{\partial A^\alpha} = -\frac{1}{c} J_\alpha$$

$$\text{or } \partial^\beta F_{\beta\alpha} = \frac{4\pi}{c} J_\alpha$$

Now the field strength tensor  $F^{\alpha\beta} = \partial^\alpha A^\beta - \partial^\beta A^\alpha$  is nothing more than the E and B fields:

$$\vec{B} = \vec{\nabla} \times \vec{A} = \epsilon_{ijk} \partial_j A_k$$

$$\vec{E} = -\vec{\nabla}\Phi - \frac{1}{c} \frac{\partial \vec{A}}{\partial t} \quad E_i = -\partial_i A_0 - \partial_0 A_i$$

$$F^{\alpha\beta} = \begin{bmatrix} 0 & -E_x & -E_y & -E_z \\ E_x & 0 & -B_z & B_y \\ E_y & B_z & 0 & -B_x \\ E_z & -B_y & B_x & 0 \end{bmatrix}$$

$$F_{\alpha\beta} = \begin{bmatrix} 0 & E_x & E_y & E_z \\ -E_x & 0 & -B_z & B_y \\ -E_y & +B_z & 0 & -B_x \\ -E_z & -B_y & +B_x & 0 \end{bmatrix}$$

so the inhomogeneous field equation is just ( $J_\alpha = (c\rho, \vec{J})$ )

$$\alpha=0 \quad \nabla \cdot \vec{E} = 4\pi \rho$$

$$\alpha=i \quad \nabla \times \vec{B} - \frac{1}{c} \frac{\partial \vec{E}}{\partial t} = \frac{4\pi}{c} \vec{J} \quad \begin{array}{l} 2 \text{ of } 4 \text{ Maxwell} \\ \text{equations!} \end{array}$$

The homogeneous equations vanish by construction

$$(\nabla \cdot \vec{B} = 0 \rightarrow \vec{B} = \nabla \times \vec{A})$$

~~Prove that~~ In covariant ~~constant~~ language they are

$$\partial^\alpha F^{\beta\gamma} + \partial^\beta F^{\gamma\alpha} + \partial^\gamma F^{\alpha\beta} = 0 \quad \alpha \neq \beta \neq \gamma$$

$$\text{or } \partial_\alpha \mathcal{F}^{\alpha\beta} = 0$$

where  $\mathcal{F}^{\alpha\beta} \equiv$  "dual field strength tensor"

$$\mathcal{F}^{\alpha\beta} = \frac{1}{2} \epsilon^{\alpha\beta\gamma\delta} F_{\gamma\delta} = \begin{bmatrix} 0 & -B_x & -B_y & -B_z \\ & 0 & E_z & -E_y \\ & & 0 & E_x \\ & & & 0 \end{bmatrix}$$

$$\partial_\alpha \mathcal{F}^{\alpha\beta} = \frac{1}{2} \epsilon^{\alpha\beta\mu\nu} \partial_\alpha F_{\mu\nu}$$

$$= \frac{1}{2} \epsilon^{\alpha\beta\mu\nu} \partial_\alpha (\partial_\mu A_\nu - \partial_\nu A_\mu)$$

$= 0$  because for each term,  $\epsilon$  is antisymmetric

While we're at it, let's think about the Lagrangian ~~and~~ Hamiltonian.

$$L = \int d^3x \mathcal{L}$$

$$\mathcal{L} = -\frac{1}{16\pi} F_{\mu\nu} F^{\mu\nu} : \text{look at } 4 \times 4 \text{ tensors, sum up}$$

$$\mathcal{L} = \frac{1}{8\pi} (E^2 - B^2)$$

Hamiltonian? Write  $\mathcal{L} = -\frac{1}{8\pi} \left[ \frac{\partial A_\mu}{\partial x^\nu} - \frac{\partial A_\nu}{\partial x^\mu} \right] \frac{\partial A^\mu}{\partial x^\nu}$

Need canonical coordinates - obvious choice the 4  $A_\mu$ 's

Canonical momenta?

$$k=1,2,3 \quad \pi^k = \frac{\partial \mathcal{L}}{\partial \dot{A}_k} = -\frac{1}{8\pi} \cdot 2 \cdot \left[ \frac{\partial A_k}{\partial x^0} - \frac{\partial A_0}{\partial x^k} \right] = \frac{E_k}{4\pi}$$

$$\pi^0 = \frac{\partial \mathcal{L}}{\partial \dot{A}_0} = 0 \Rightarrow \mathcal{H} = \sum_{k=1}^3 \pi^k \dot{A}_k - \mathcal{L} = \frac{1}{4\pi} \left( \frac{1}{2} (E^2 + B^2) + \vec{E} \cdot \vec{\nabla} \Phi \right)$$

At this point things get a bit tricky.

Note (first of all) that there is no momentum conjugate to  $A_0$ :  $A_0$  has no independent dynamics

Second, in the absence of sources, Maxwell's equations (which we could as well derive as a set of Hamilton equations) require

$$\vec{\nabla} \cdot \vec{E} = 0$$

so the 3 E's are not independent

The system has constraints. They are related to gauge symmetry; not all the A's are independent variables.

To the best of my knowledge, this is not a problem for the classical theory, but it is a problem in constructing the quantum theory using "canonical quantization":

- 1) write down classical  $\mathcal{L}$
- 2) construct classical  $H$  & identify  $p, q$ .
- 3) Impose canonical quantization conditions

$$[q_i, p_j] = i\hbar\delta_{ij}$$

- 4) Construct ~~classical~~ quantum  $H$  by replacing coordinates & momenta by coord & mom. operators.

to have  $\rightarrow$   
energy  
expansion

To proceed beyond point 3) it is necessary to fix a gauge ~~to have explicit energy~~ and in order to make  $H$  time independent the gauge must be non covariant, like Coulomb gauge. But then the formulas all look non covariant (there is an instantaneous Coulomb interaction, for example) and then it is a lot of work to show that the quantum theory is covariant - all the non covariant parts cancel.



Anyway, back to Lagrangians. The  $A_\mu$ 's form a four vector and transform under Lorentz transformations like any 4-vector. In matrix notation where the L.T. is

$$\Lambda = \exp(i\omega \cdot \mathbf{J} - \mathbf{z} \cdot \mathbf{K})$$

$$\begin{pmatrix} A_0' \\ \vec{A}' \end{pmatrix} = \Lambda \begin{pmatrix} A_0 \\ \vec{A} \end{pmatrix}$$

$E$  and  $B$  are elements of the second rank tensor  $F^{\mu\nu}$ . It transforms as

$$F'^{\alpha\beta} = \frac{\partial x'^\alpha}{\partial x^\mu} \frac{\partial x'^\beta}{\partial x^\nu} F^{\mu\nu}$$

or  $F' = \bar{\Lambda} F \Lambda^T$  (matrix notation)

Let's multiply out for a boost along the  $x$  axis:

$$\Lambda = \begin{bmatrix} \gamma & +\beta\gamma & 0 & 0 \\ +\beta\gamma & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

$$\begin{aligned} & \frac{\gamma - \gamma^2 \beta^2}{\gamma + 1} \\ & = \gamma - \frac{\gamma^2 \beta^2}{\gamma + 1} \\ & = \frac{\gamma^2 + \gamma - \gamma^2 \beta^2}{\gamma + 1} \\ & = \frac{\gamma^2(1 - \beta^2) + \gamma}{1 + \beta} = \frac{1 + \gamma}{1 + \beta} \end{aligned}$$

General result

$$\begin{aligned} E_1' &= E_1 & B_1' &= B_1 \\ E_2' &= \gamma(E_2 - \beta B_3) & B_2' &= \gamma(B_2 + \beta E_3) \\ E_3' &= \gamma(E_3 + \beta B_1) & B_3' &= \gamma(B_3 - \beta E_2) \end{aligned}$$

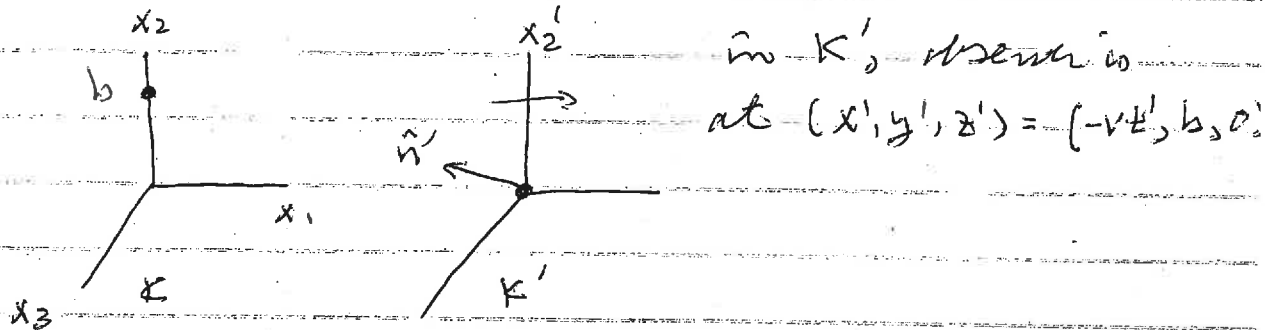
This I checked

$$\begin{aligned} \vec{E}' &= \gamma(\vec{E} + \vec{\beta} \times \vec{B}) - \frac{\gamma^2}{\gamma + 1} \vec{\beta}(\vec{\beta} \cdot \vec{E}) \\ \vec{B}' &= \gamma(\vec{B} - \vec{\beta} \times \vec{E}) - \frac{\gamma^2}{\gamma + 1} \vec{\beta}(\vec{\beta} \cdot \vec{B}) \end{aligned}$$

← This I did not.

Example: point charge at origin in frame  $K'$ , point charge moving with  $\vec{v} = +\hat{x}v$  in frame  $K$

In  $K$ , observer is at  $(x, y, z) = (0, b, 0)$  but



in  $K'$ , observer is at  $(x', y', z') = (-vt', b, 0)$

Also in  $K'$ ,  $\vec{B}' = 0$ ,  $\vec{E}' = \frac{q \hat{r}'}{(r')^2}$

or  $E_1' = -\frac{qv t'}{(r')^3}$      $E_2' = -\frac{qb}{(r')^3}$      $E_3' = 0$

and  $(r')^2 = b^2 + v^2 t'^2$

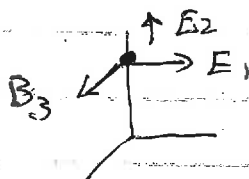
We have to convert both the coordinates and the fields. To convert the coords,  $t' = \gamma t$  so

$$E_1' = \frac{q(-\gamma v t)}{[b^2 + (\gamma v t)^2]^{3/2}} \quad E_2' = \frac{q b}{[b^2 + (\gamma v t)^2]^{3/2}}$$

Now we transform the fields

$$E_1 = E_1' = -\frac{q \gamma v t}{[b^2 + (\gamma v t)^2]^{3/2}}, \quad E_2 = \frac{q b \gamma}{[b^2 + (\gamma v t)^2]^{3/2}} (= \gamma E_2')$$

$$B_3 = \beta \gamma E_2' = \beta E_2$$

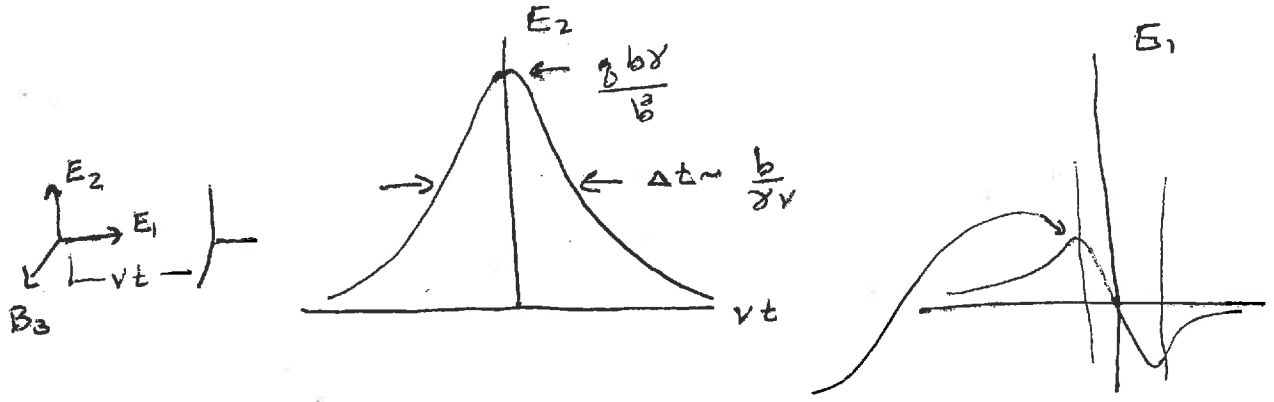


Note transverse B ( $B_3$ ) and as  $v \rightarrow c$   $|B_3| \approx |E_2|$

Note NR limit:  $\vec{B} = \frac{q}{c} \frac{\vec{v} \times \vec{r}}{r^3}$  ( $\frac{v}{c}$  is CGS)

(Biot-Savart:  $q\vec{v} \rightarrow I d\vec{l}$ )  $\frac{q}{c} \frac{(\vec{v} \times \vec{r})}{r^3}$

Note extreme relativistic limit  $E_2 \propto \gamma$  but  $E_1 \sim 1$



at  $\gamma vt = b$   $E_1 = \frac{qb}{b^3 2^{3/2}} \sim \frac{b}{b^2}$

Fields become a pulse of plane wave



$q$  at rest

indeed if detector averages over time  $T > \frac{b}{\gamma v}$ ,  $\langle E_1 \rangle = 0!$

[Future: can exploit analogy between fields of rel particles and plane wave  $\equiv$  "Weisaker-Williams approximation"]

Why  $F_{\mu\nu}^2$ ?

So why Maxwell's equations? Why is

$$\mathcal{L} = -\frac{1}{4\pi} F_{\mu\nu} F^{\mu\nu} + \text{nothing else?}$$

In pure classical case, only justification is observation of superposition

The answer involves dynamics (and the answer is, really, there is something else!)

Let's work in units with  $\hbar = c = 1$ , so in dimensions <sup>+ mass = energy</sup>

$$[\text{energy}] = \frac{1}{[\text{length}]} \quad (\hbar c = \text{energy} \times \text{length}).$$

A Lagrange density  $\mathcal{L}$  is  $\frac{[\text{energy}]}{[\text{length}^3]} = \frac{1}{[\text{length}]^4}$

$F_{\mu\nu} \sim \vec{E}$  ( $\sim g/v^2$ ) ~~is~~ and  $g$  is dimensionless ( $\frac{q^2}{\hbar c} = \frac{1}{137}$ )  
so  $[\vec{E}] \sim \frac{1}{[\text{length}]^2}$  and so

$$\mathcal{L} = (\text{dimensionless \#}) \cdot \frac{1}{([\text{length}]^4)} F_{\mu\nu} F^{\mu\nu} + \dots$$

What about other terms? Gauge invariance + Lorentz invariance require  $\mathcal{L} = \mathcal{L}(F_{\mu\nu} F^{\mu\nu})$  or  $F\tilde{F}$

and so if we imagine  $\mathcal{L}$  is a polynomial in  $F^2$ , each new term must be ~~multiplied~~ by a parameter of dimensions  $[\text{length}]^4$  or  $\frac{1}{[\text{some mass}]^4} = \frac{1}{\Lambda^4}$

What is the mass? If we are thinking about ~~some~~ electrodynamics the "mass" ~~is~~ is the scale at which new physics modifies Maxwell's eqns. The simplest new physics is the fundamentally quantum process of the exchange of virtual particles as intermediate

states in ~~one process~~ (production of  $e^+e^-$  pairs) and so  $\Lambda \sim m_e$ . Now in ~~one~~ wave numbers / momentum

space,  $F^2 \sim k^2$ ,  $\frac{F^4}{\Lambda^4} \sim \frac{k^4}{\Lambda^4}$  (notice this is even true classically:  $k \sim \Lambda$ , so nothing happens...)

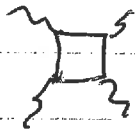
$$F = \partial_\mu A_\nu - \partial_\nu A_\mu \approx k_\mu A_\nu - k_\nu A_\mu$$

so for "typical momenta"  $k \ll \Lambda$  ( $= m_e \sim \frac{1}{2} \text{ MeV}$ ) such processes are completely negligible. And of course at  $k \sim m_e$  these new processes turn on, and a classical description ~~is~~ fails completely (of course it has already failed for below that - in the actual radiation processes in atoms). Indeed a calculation to support this statement was done by 2 of Heisenberg's students, Euler and Kockel, in 1935 - they started with QED (photons + electrons) and integrated <sup>out</sup> the electrons ~~and~~ to derive an effective theory of photons

$$\mathcal{L} = \frac{1}{2} (E^2 - B^2) + \frac{e^4}{360\pi^2 m_e^4} \left[ (E^2 - B^2)^2 + 7(E \cdot B)^2 \right] + \dots$$

valid for frequencies  $\hbar\omega \ll m_e c^2$ . (So not even  $\frac{1}{m_e^4} \gg \frac{d^2}{m_e^4}$  which is  $10^{-4}$  as big!)

The new term includes "photon splitting"  $\gamma \rightarrow \gamma\gamma$  interactions as explicit nonlinear terms in  $\mathcal{L}$ .



QED is itself an incomplete ~~the~~ description of Nature at very high energy, too... it is only a part of a combined ~~interaction~~ "electroweak" interaction.